X-ray Photoelectron Spectroscopy (XPS)

As part of the course 'Characterization of Catalysts and Surfaces'

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Resource for further reading: *Surface Analysis by Auger and X-ray Photoelectron Spectroscopy*, D. Briggs, J.T. Grant, eds., IM Publications and SurfaceSpectra Ltd., 2003

XPS in a nut-shell

- X-ray photoelectron spectroscopy (XPS) is a classical method for the semiquantitative analysis of surface composition
- It is also referred to as Electron Spectroscopy for Chemical Analysis (ESCA)
- It is based on the photoelectric effect, i.e., emission of electron following excitation of core level electrons by photons
- It is surface sensitive because of the low inelastic mean free path of electrons
- An XPS setup consists of a X-ray source, a sample chamber and an electron analyzer
- XPS requires a monochromatic source of X-rays, i.e., either from a lab based anode or from a synchrotron, with a X-ray monochromator in both cases
- Traditionally, XPS works only in ultrahigh vacuum because of scattering of electrons in gases
- Since recently, XPS can also be performed in the mbar pressure range

This is a photoelectron spectrum of TiO₂



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This is a lab based XPS instrument



Related methods

Method	Particle in	Energy range (eV)	Particle out	Information, electronic level	Technique
Photoemission	Photon	100-10000	Electron	Filled core states	XPS
Photoemission	Photon	0-10 (100)	Electron	Filled valence states	UPS
Inverse photoemission	Electron	100-1000	Photon	Empty states	IPES
Electron energy loss	Electron	10-10000	Electron	Electronic & vibrational transitions	(HR)EELS
Auger	Electron or Photon	10-10000	Electron	Filled states	AES
Absorption/emi ssion	Photon	10-2000	Photon	Electronic transitions, filled states	XRF, XAS, RIXS

The photoemission process



 $KE = hv - BE - \Phi$ for a solidKE = hv - IP for a gas Φ : photoelectric workfunction (4-6 eV)

The binding energy



Fate of core hole



Core hole relaxation: Auger's for low Z – XRF for high Z



A photoelectron spectrum in more detail



Photoemission Intensity (arb. units)

Electron inelastic mean free path



XPS probe depth



Contribution of atom in depth d to PE peak:

$$I = I_0 \exp\!\!\left(\frac{-d}{\lambda\cos\theta}\right)$$

- Only electrons from close to surface can escape without energy loss
- Inelastic mean free path is a strong function of kinetic energy.
- Exponentially decreasing contribution from atoms below the surface to the photoelectron peaks (without loss)
- At close to normal detection angles 95% of the signal comes from atoms within 3λ of the surface
- Molecular scale surface sensitivity
- > Probing depth varies with kinetic energy.
- Depth profiles can be obtained by varying the incident photon energy with a tunable X-ray source, i.e., at a synchrotron, or by varying the detection angle

Inelastic background



- Electrons from deeper below the surface loose energy and emerge with reduced KE (increased apparent BE)
- Electrons very deep in surface loose all energy and cannot escape
- XPS spectra show characteristic "stepped" background (intensity of background towards higher BE of photoemission peak is always greater than towards lower BE)







- n: principal quantum number
 - I: orbital angular momentum quantum number
 - s: spin angular momentum quantum number
 - j = || ± s|: total angular momentum quantum number
- For I = 0, s levels are singlets, no splitting
- For I > 0, p,d,f levels give rise to doublets, because the unpaired electron left in an orbital can have its spin and orbital angular momentum either parallel or anti-parallel
- The degeneracy, 2j+1, determines how many possibilities exist for parallel or anti-parallel pairing and thus the relative peak ratio of the two doublet components.
- BE of lower j value in doublet is higher (BE 2p1/2 > BE 2p3/2)
- Magnitude of spin-orbit splitting increases with Z
- Magnitude of spin-orbit splitting decreases with distance from nucleus (increased nuclear shielding)







Auger peaks

- Result from excess energy of atom during relaxation (after core hole) creation
- always accompany XPS
- broader and more complex structure than photoemission peaks, because of different valence levels involved and secondary effects
- Nomenclature: ABC, with A: level of initial core hole, B: level from which A is filled, C: level from which Auger electron is emitted. Example: KLL
- KE of Auger electrons are independent of incident photon energy: KE_{ABC} = BE_A - BE_B -BE_C*.

 ${\rm BE_{c}}^{*}$ is slightly different from ${\rm BE_{c}}$ because of the hole in level B

Auger spectra are sensitive to chemistry, because of the involvement of outer shells.

Core level chemical shifts

- Position of orbitals in atom is sensitive to chemical environment of atom
- In gas phase XPS, the differences in core electron ionization energies are directly apparent
- In solids, all core levels for that atom shifted by approx. same amount (<10 eV)</p>
- Chemical shift correlated with overall charge on atom (Reduced charge increased BE)
 - number of substituents
 - substituent electronegativity
 - formal oxidation state (unreliable, depending upon ionicity/covalency of bonding)
- Chemical shift analysis is powerful tool for chemical composition, functional group and oxidation state analysis



Further features

Final state effects

- Screening and relaxation effects: satellite peaks on low or high BE side
- Important in transition metal oxides
- Useful in fingerprinting of complex systems
- Descrete electron structures due to electronic excitations, e.g., (surface) plasmons
- Asymmetric peak shapes in metals due to low energy electron-hole excitations



FIG. 1. Cu 2p photoelectron spectra of Cu dihalides. The ines leading to a final state with a ligand hole (L) show a chemical shift.





Figure 2. The XPS 3d core level spectra of metallic Mo, MoO_2 and MoO_3 .

Werfel and Minni, 1983



Electron energy analyzer

- Circular orbit of electrons from entrance slit to detector, controlled by inner and outer potential
- Since the resolution, ΔE/E, depends on kinetic energy, an electrostatic lens is used to retard electrons to constant kinetic energy, referred to as pass energy, E₀

 $\frac{\Delta E}{E_0} = \frac{s}{2R_0}$

- s slit width
- R₀ radius of electron orbit through hemisphere



Hemispherical electron energy analyser

Estimation of signal intensities

Contribution of element A in depth d to photoemission signal



- Cross sections range up to about 10⁻¹⁸ cm²
- 10¹⁵ atoms cm⁻² (equivalent to a monolayer) lead to about 10⁻³ photoelectrons per incident photon
- ➤ Typical photon flux: 10¹² s⁻¹, leads to about 10⁹ photoelectrons s⁻¹
- > For most elements, sensitivity is in the range of 0.1 to 1 % of a monolayer

Photoionisation cross sections for Ti core levels



J.J. Yeh, Atomic Calculation of Photoionization Cross-Sections and Asymmetry Parameters, Gordon and Breach Science Publishers, Langhorne, PE (USA), 1993, and J.J. Yeh and I.Lindau, Atomic Data and Nuclear Data Tables, **32**, 1-155 (1985).

How to get the binding energy scale



varying KE by changing photon energy



Depth profile



Klyushin et al., 2014

Auger yield NEXAFS

- Excitation into unoccupied molecular orbital (LUMO)
- Resonant absorption



- Auger electron emission following filling of core level
- Auger electron emission intensity scales with rate of initial excitation into LUMO
- Surface sensitive NEXAFS technique



Auger yield NEXAFS



High pressure XPS setup



X-rays

MeOH oxidation on Cu catalyst



In situ gas and surface characterisation under catalyst operating conditions (400°C, few mbar MeOH, Bluhm et al., 2004)

Normalized count rate (a.u.)



Identification of sub-surface oxygen species as decisive for catalyst activity

Bluhm et al., JPCB 2004

Other XPS developments



Nanoparticles in liquids



Electrocatalysis



El Gabaly et al. PCCP (2010) ³¹

Brown et al., Langmuir (2013)