

Last time: HHG with a semiclassical model

1st step: Quantum \Rightarrow tunneling

2nd step: classical \Rightarrow propagation in laser field

3rd step: Quantum \Rightarrow recombination

showed that dominant contribution to HHG comes from electrons that:

- i) Return to the nucleus
- ii) Appear in the continuum with zero velocity
- iii) Have an appropriate KE to produce a given harmonic @ the time of return: $KE = (2n+1)\hbar - I_p$

Fully Quantum and exact theory ^{of HHG} has been formulated

i) as a solution to TDSE (early '90s)

ii) or, equivalently in terms of solutions of time-indep. Floquet equations

Fully quantum approach: TDSE not so easy to solve, esp. for elliptical polarizations. Gives little physical insight

On the other hand: Semiclassical ("simple man's model")

\Rightarrow good physical insight

\Rightarrow does not account for quantum diffusion/interference

Today: present an approximate analytical solution to the Schrodinger eqn. as an intermediary btwn the semiclassical & the exact Quantum treatment

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First: a simple example to see how radiation is produced when the electron transitions between different quantum states \Rightarrow Need dipole moment $\langle x \rangle$

Wavefunction of electron in a state of quantum number n & Energy E_n : $\psi_n = \psi_n e^{-iE_n t}$
 $\psi_n^* = \psi_n^* e^{+iE_n t}$
time-indep. part

Expectation value $\langle x \rangle$ of position

of such an electron: $\langle x \rangle = \int_{-\infty}^{\infty} x \psi_n^* \psi_n dx$
x-component of dipole moment

$\langle x \rangle = \int_{-\infty}^{\infty} x \psi_n \psi_n^* dx \Rightarrow$ constant in time!

\Rightarrow electron does not oscillate & no radiation occurs

\Rightarrow an electron in a specific quantum state produces no radiation.

Now, suppose the electron wave-function is a superposition of 2 states: $\psi = a \psi_n + b \psi_m$

$|a|^2 \Rightarrow$ probability the electron is in state a

$|b|^2 \Rightarrow$ " " " " " " b

$\langle x \rangle = \int_{-\infty}^{\infty} x (a^* \psi_n^* + b^* \psi_m^*) (a \psi_n + b \psi_m) dx$

$= \int_{-\infty}^{\infty} x (\underbrace{a^2 \psi_n^* \psi_n}_{\text{constant}} + \underbrace{b^* a \psi_m^* \psi_n + a^* b \psi_n^* \psi_m}_{\text{time-dep. terms}} + \underbrace{b^2 \psi_m^* \psi_m}_{\text{constant}}) dx$

time-dep. part: $\underbrace{b^* a \int_{-\infty}^{\infty} x \Psi_m^* \Psi_n dx}_{\text{time-indep}} \cdot e^{it(E_m - E_n)}$

$+ a^* b \int_{-\infty}^{\infty} x \Psi_n^* \Psi_m dx \cdot e^{-it(E_m - E_n)}$

Oscillation @ frequency: $\omega = \frac{E_m - E_n}{\hbar} \Rightarrow$ frequency of the emitted/absorbed photon is a $m \leftrightarrow n$ transition!

atomic dipole: $d_x = \int_{-\infty}^{\infty} x \Psi_m^* \Psi_n dx = \langle \Psi_m | x | \Psi_n \rangle$
 for $n \rightarrow m$ transition

transitions where $d_x = 0$ are forbidden!

Hence transitions where $\int_{-\infty}^{\infty} x \Psi_n^* \Psi_m dx \neq 0$ are allowed

Ψ_n & Ψ_m can be calculated exactly for a hydrogenic atom, where the wavefunction is dictated

by the quantum #'s n, l, m_l
 energy \uparrow orbital ang. momentum \uparrow magnetic quantum #

condition $\int_{-\infty}^{\infty} \vec{r} \Psi_{n,l,m_l}^* \Psi_{n',l',m_l'} dV \neq 0$

determines allowed transitions \rightarrow "dipole selection rules"

Selection rules: $\Delta l = \pm 1$; $\Delta m_l = 0, \pm 1$

Note: These rules result in an emission/absorption of a single photon of energy

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$|E_m - E_n|$, which carries off angular momentum $\pm \hbar$ (equal to the difference in angular momentum, l , between m & n states)

\Rightarrow so to get radiation frequencies in HHG, we have to calculate the x-component of time-dep. dipole moment:

$$x(t) = \langle \Psi(t) | x | \Psi(t) \rangle$$

why only $x(t)$? Because electron osc. involved in HHG is along the \hat{x} -axis, since $\vec{E} = E \cos(\omega t) \hat{x}$

\Rightarrow Approximate solution for $|\Psi(t)\rangle$ already obtained from SFA.

\Rightarrow Note: $|\Psi(t)\rangle$ is made up of a continuum of states $|v\rangle$ + ground state $|0\rangle$:

$$|\Psi(t)\rangle = e^{iE_0 t} a(t) |0\rangle + \int d^3v \underbrace{a_v(t)}_{\text{integration over a continuum of velocity states}} |v\rangle$$

i) In case of no depletion: $a(t) \approx 1$

ii) Note: only ground state & continuum states contribute to $|\Psi(t)\rangle \Rightarrow$ no intermediate resonances.

iii) Neglect Coulomb field in the continuum (key SFA assumption)

where $a_v(t) = \langle v | \Psi(t) \rangle$ was derived

previously using SFA for the following TDSE:

$$i \frac{\partial}{\partial t} |\Psi\rangle = \left[\underbrace{-\frac{1}{2} \nabla^2 + V(x)}_{\hat{H}_0} - \underbrace{E \cos(\omega t) x}_{\text{length gauge: } \hat{V}_L} \right] |\Psi\rangle$$

\hat{H}_0 (field free Hamiltonian): $-\frac{1}{2} \nabla^2 + V(x)$ \leftarrow atomic potential

\hat{V}_L (interaction with the laser field): $\hat{V}_L = \underbrace{E \cos(t)}_{\text{field strength}} \cdot \underbrace{\hat{x}}_{\text{dipole term}}$

Full solution of TDSE:

$$|\Psi(t)\rangle = i \int_0^t dt' e^{-i \int_{t'}^t \hat{H}(t'') dt''} \hat{V}_L(t') e^{-i \int_0^{t'} \hat{H}_0(t'') dt'' + i I_p t'} |0\rangle + e^{+i I_p t} |0\rangle$$

SFA: set $\hat{H} \approx \hat{H}_0 + \hat{H}_{\text{Laser}} \approx \hat{H}_{\text{Laser}} = \frac{1}{2} m \hat{V}^2 \Rightarrow$ KE of a free electron

$$a_v(t) = \langle v | \Psi(t) \rangle \approx i \int_{t_0}^t dt' e^{-i \int_{t'}^t \frac{(V'')^2}{2} dt'' + i I_p t'} \cdot \underbrace{\langle v(t') | x | 0 \rangle}_{\text{dipole matrix element!}} \times E \cos(t')$$

From conservation of canonical momentum: $\vec{p} = \vec{v}(t) + \vec{A}(t) = \vec{v}(t') + \vec{A}(t')$

$$\left(\frac{V''}{2}\right)^2 = \left[\frac{\vec{v}(t) + \vec{A}(t) - \vec{A}(t')}{2}\right]^2$$

Dipole moment: $x(t) = \langle \Psi(t) | x | \Psi(t) \rangle$

$$x(t) = \langle 0 | e^{-i I_p t} + \int d^3 v' a_{v'}^*(t) \langle v' | x | v \rangle (e^{i I_p t} | 0 \rangle + \int d^3 v a_v(t) | v \rangle$$

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again, only cross-terms $\langle 0|V\rangle$ & $\langle V|0\rangle$ contribute to the radiation

Note: continuum - continuum contributions: $\langle v'|V\rangle$ are neglected \Rightarrow one reason is that ionization probability is low (no saturation), hence $a_v^* a_v \ll a_v \Rightarrow$ these terms only introduce higher order corrections \Rightarrow corresponds with the intuition that the main contribution to HHG come from recombination with the ground state, $|0\rangle$

$$X(t) \approx \int dV^3 a_v(t) \cdot \underbrace{\langle 0|X|V\rangle}_{d_x^*(t)} + \int dV^3 a_v^*(t) \underbrace{\langle V|X|0\rangle}_{d_x(t)}$$

$$= 2 \cdot \text{Re} \left(\int dV^3 a_v(t) \cdot d_x^*(t) \right)$$

substituting for $a_v(t)$:

$$X(t) = 2 \text{Re} \left\{ i \int_0^t dt' \int d^3p \cdot E \cos(t') \underbrace{\langle v(t')|X|0\rangle}_{d_x(t')} \times d_x^*(t) \times e^{-iS(\vec{p}, t, t')} \right\}$$

$$\text{where } S(\vec{p}, t, t') = \int_{t'}^t dt'' \frac{[\vec{p} - \dot{A}(t'')]^2}{2} - I_p t'$$

Note: t' corresponds to ionization time

As we learned in the last lecture, there is 1 particular trajectory for each ionization time, t' , that contributes

⇒ can use the saddle-point method to evaluate the integral $\int d^3p \dots$

$$S = \int p dq$$

$$\text{SPM: } \nabla_{\vec{p}} S(\vec{p}, t, t') = 0$$

Note $\nabla_p S(\vec{p}, t, t') = \vec{x}(t) - \vec{x}(t')$ ⇒ hence

$\nabla_{\vec{p}} S(\vec{p}, t, t')$ corresponds to requiring that electron that appears in the continuum @ time t' returns to the same pt @ time t .

$$S(\vec{p}, t, t') = \frac{1}{2} \int_{t'}^t dt'' \left\{ (P_x - \underbrace{E \sin(t'')}_{A_x})^2 + P_y^2 + P_z^2 \right\} - I_p t'$$

$$\frac{\partial S}{\partial P_y} = \frac{\partial S}{\partial P_z} = 0 \Rightarrow P_y = P_z = 0 \quad \Rightarrow \text{justifies the semi-classical assumption of zero transverse velocity}$$

$$\frac{\partial S}{\partial P_x} = \int_{t'}^t dt'' (P_x - E \sin(t'')) = P_x(t-t') - E \cos(t'') \Big|_{t'}$$

$$P_{st}(t, t') = \frac{E(\cos(t) - \cos(t'))}{t-t'} \quad \Rightarrow \text{corresponds to classical trajectory}$$

→ trajectory that contributes to the harmonic emission for ionization & recombination times given by t' & t , respectively.

$t-t'$ ⇒ excursion time
long trajectories, smaller P_{st} , ionization near the peak