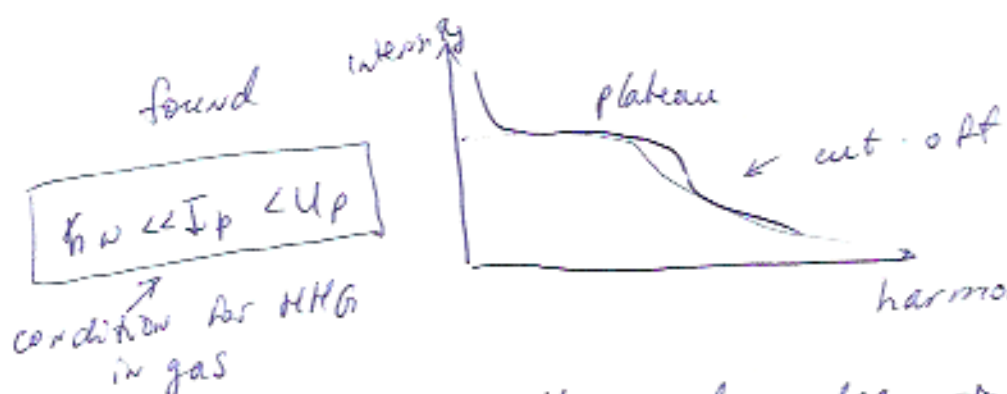


Last time: HHG using atoms in a gas phase



Helium: $I_p = 0.9 \text{ au}$
 $U_p \text{ (for } F=0.1 \text{ au)} = 1.7 \text{ au}$
 $E_{\text{max}} = h\omega_{\text{max}} \approx I_p + 3.17U_p = 6.289$
 @ $800 \text{ nm } (\omega=0.057) \Rightarrow n=110$

\Rightarrow make up an attosecond pulse \Rightarrow can get harmonics in the XUV range $40 - 120 \text{ nm}$

ionizing radiation of Ti: sapphire $\sim 800 \text{ nm}$
 \Rightarrow cut-off wavelength λ (for $n=110$): $\lambda_{\text{cut-off}} \approx 7 \text{ nm}$ (X-ray)

Today \Rightarrow HHG & attosecond pulse generation on plasma mirrors

\Rightarrow higher frequency (in the X-ray range):
 $\lambda = 0.01 - 10 \text{ nm} \Rightarrow$ X-ray range

\Rightarrow inner shell dynamics of atoms (characteristic of transition freq. from ~~inner~~ inner shells).

\Rightarrow higher energy flux

\Rightarrow as pump - as probe experiments
 current experiments use as a probe & infrared as a pump or vice-versa.

Example: tunnel ionize an atom with an ultra-short Ti: sapphire laser pulse: FWHM $\sim 5 \text{ fs}$,
 $\lambda \approx 800 \text{ nm}$; then probe the evolution of a hole left behind by a transient absorption of an as pulse
 Goulielmos, Nature 466, 739, 2010

Plasmas \Rightarrow can go to much higher laser intensity \Rightarrow instead of $10^{15} \frac{W}{cm^2} \Rightarrow 10^{22} \frac{W}{cm^2}!$

(ultra-relativistic regime)

\Rightarrow solid surface, but turns to plasma due to high laser intensity

\Rightarrow First experimental demonstration of using plasma mirrors to generate as pulses \Rightarrow

Nomura, et al, Nature Phys. 5, 124, 2009

First, why can't go above $10^{15} \frac{W}{cm^2}$ in ^{gases?} ~~gases~~.

\Rightarrow depletion due to high intensities

\Rightarrow poor phase matching

\Rightarrow magnetic fields become important \Rightarrow introduce drifts

(no recollision step because of B fields!)

HHG in plasmas \Rightarrow focus intense, ultrashort pulse on

a solid target

\Rightarrow due to time-scales (fs), very little plasma expansion

\Rightarrow laser pulse reflects on dense plasma

\Rightarrow this is called plasma mirror

Two mechanisms behind HHG on plasma mirrors:

1) Coherent Wake emission (CWE)

2) Relativistic Oscillating mirror (ROM)

} very different harmonic spectra

3 steps of HHG using CWF

(i) Electrons @ the plasma surface are pulled into a vacuum & then slammed back into the plasma by the laser field

(ii) Electrons propagate in the dense part of the plasma, forming ultra-short bunches, exciting local plasma oscillations

\Rightarrow frequency of plasma osc. depend on depth
deeper \Rightarrow higher frequency

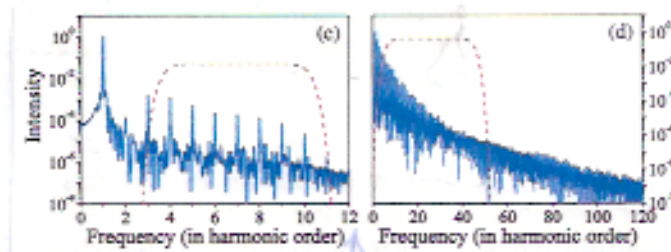
(iii) high harmonics are generated with frequency given by the local plasma freq.

\Rightarrow cut-off given by the maximum plasma frequency, $\omega_p^{max} \approx n \cdot \omega_{Laser}$ where $n \approx 15-30$

\Rightarrow ROM is the dominant contribution beyond ω_p^{max}

ROM HHG spectra

CWE
HHG
spectra



cut-off given by ω_p^{max}

much higher cut-off in ROM compared to CWE

When do different regimes contribute?

(4)

CWE vs ROM?

Parameter a_0 that depends on intensity & frequency of laser: $a_0 = \frac{eE}{\hbar\omega} = \frac{eE}{m\omega c}$
normalized vector potential

For Intensity = $10^{15} \frac{W}{cm^2}$; 800 nm lld

$$a_0 \approx 0.03$$

CWE dominates the harmonic signal for $a_0 \lesssim 1$

\Rightarrow need intensity $\sim 10^{18} \frac{W}{cm^2}$ for ROM to

become significant

\Rightarrow therefore need to go to intensity $> 10^{18} \frac{W}{cm^2}$ to get harmonics in the x-ray regime (since CWE can only go up to $n \approx 30$)

\Rightarrow ROM produced by the Doppler shift

created by the relativistically oscillating mirror surface

\Rightarrow ROM extensively studied analytically

\Rightarrow CWE only recently identified

\Rightarrow however as pulse generation achieved predominantly

using CWE in Nomura, Nat. Phys. 5, 124, 2009
 $\sim 10^{19} \frac{W}{cm^2}$

\Rightarrow HHG has been achieved with both ROM & CWE

\Rightarrow Note HHG not enough to create an as pulse \Rightarrow

need phase coherence (phase locking between different harmonics)

Nomura et al, Nat. Phys. 5, 124, 2009

How did they know CWF was the main contributor to HHG? regime $a_0 \sim 1 \Rightarrow$ intensity $\approx 4 \times 10^{18} \frac{W}{cm^2}$ (average)

- 1) Appearance of a distinct cut-off frequency (in the $n=15-30$ range) $\left. \begin{array}{l} a_0 \approx 1.5 \text{ (both ROM \& CWF can contribute)} \end{array} \right\}$
- 2) Scaling of this cut-off with target density \Rightarrow higher cut-off with higher density

- \Rightarrow ROM theoretically well studied but as pulses have not been experimentally demonstrated
- \Rightarrow CWF only recently discovered, but experimentally shown to produce as pulses
- \Rightarrow would like as pulses with ROM (much higher $\omega_{cut-off}$)
- \Rightarrow clearly still much to do for experimentalists!
- \Rightarrow For theory \Rightarrow better understanding of CWF is desirable

CWF \Rightarrow (harmonic order given by local plasma density is same as electron density, $n_e(x)$ (neutral plasma) \Rightarrow plasma freq. of osc. $n_e(x) \Rightarrow$ plasma density) \Rightarrow so can have different harmonics if there is a density gradient: $n_e(x)$, where x is the depth of the plasma

Plasma Oscillations \Rightarrow = Langmuir waves \Rightarrow rapid osc. of electron density in conducting media
 In cold neutral plasma, displacing the electrons with respect to ions, creates oscillations @ plasma frequency:

$\frac{1}{4\pi\epsilon_0} = 1$ (in a.u.)

$\omega_{pe} = \sqrt{\frac{n_e(x) e^2}{m \epsilon_0}}$

$\omega_{pe} \sim O(1) \omega_L$ and, since

$$\omega_{cut-off} = n \cdot \omega_L \quad \text{where } n = 15-30$$

$\& \omega_L \approx 0.057$

\Rightarrow so plasma freq. is close to freq. of osc. inside the hydrogen atom.

How to create a plasma mirror? what is a plasma mirror?

\Rightarrow start with a solid target

\Rightarrow plasma created at the front of the pulse,

intensities $10^{14} - 10^{16} \frac{W}{cm^2}$

\Rightarrow plasma hardly has enough time to expand
for an ultrashort pulse (under 100 fs)

\Rightarrow sharp plasma-vacuum interface

\Rightarrow the surface of low z target (such as aluminium, $z=13$)
is almost fully ionized

\Rightarrow high density $n(x) \gg n_c \leftarrow$ critical density

where n_c corresponds to the density that would
produce osc. @ frequency ω_L : $n_c = m \epsilon_0 \omega_L^2 / e^2$

material (plastic): $n = 200 n_c$; aluminium: $n = 400 n_c$
silica

Therefore, laser field is not absorbed \Rightarrow incident
light reflected \Rightarrow plasma mirror.

\Rightarrow plasma-vacuum interface is flat (due to limited expansion)

\Rightarrow pulse is not distorted on reflection