



The interplay of vertical velocity, ice microphysics, and radiative heating

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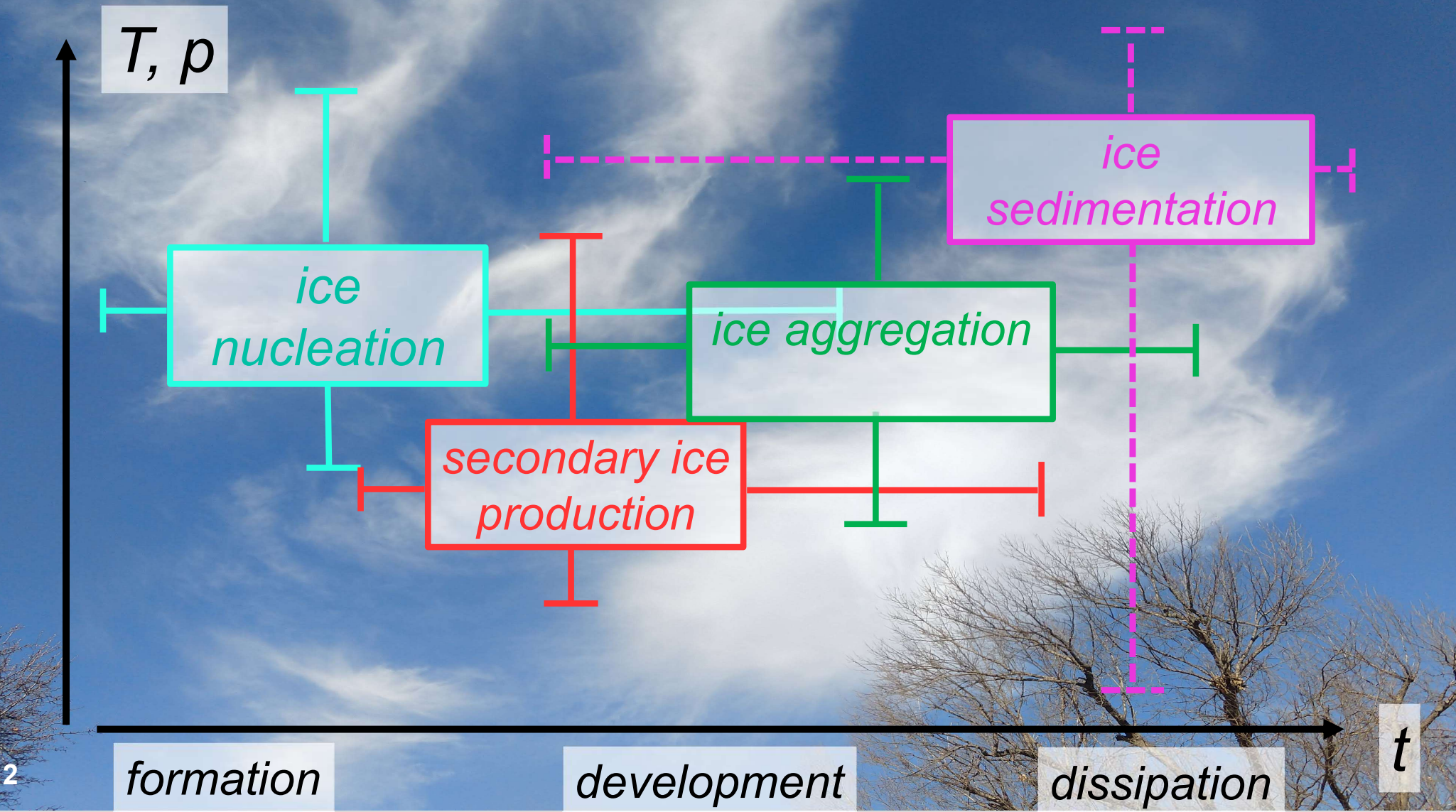
Christian Rolf and Martin Krämer, Forschungszentrum Jülich

Athanasios Nenes, EPFL

Dongmin Lee and Lazaros Oreopoulos,
NASA Goddard Space Flight Center

Tropospheric Division, Institute for Meteorology and Climate Research





2

formation

development

dissipation

t

Vertical velocity modulates and is determined by these processes.

$T,$
 p

$$|v_z| \propto \frac{dp}{dt}, \frac{dT}{dt}$$

ice nucleation

$$\frac{ds_i}{dt}, H_{sub}, H_{rad}$$

aggregation

$$K \propto a_1, c_1, a_2, c_2$$

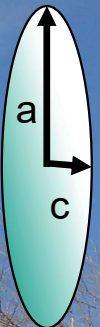
sedimentation

$$\frac{dm_i}{dt} \propto \mathcal{D}, \rho_v$$

$$v_t \propto m_i$$

secondary production

$$K \propto v_{t,1}(m_1), v_{t,2}(m_2)$$



t

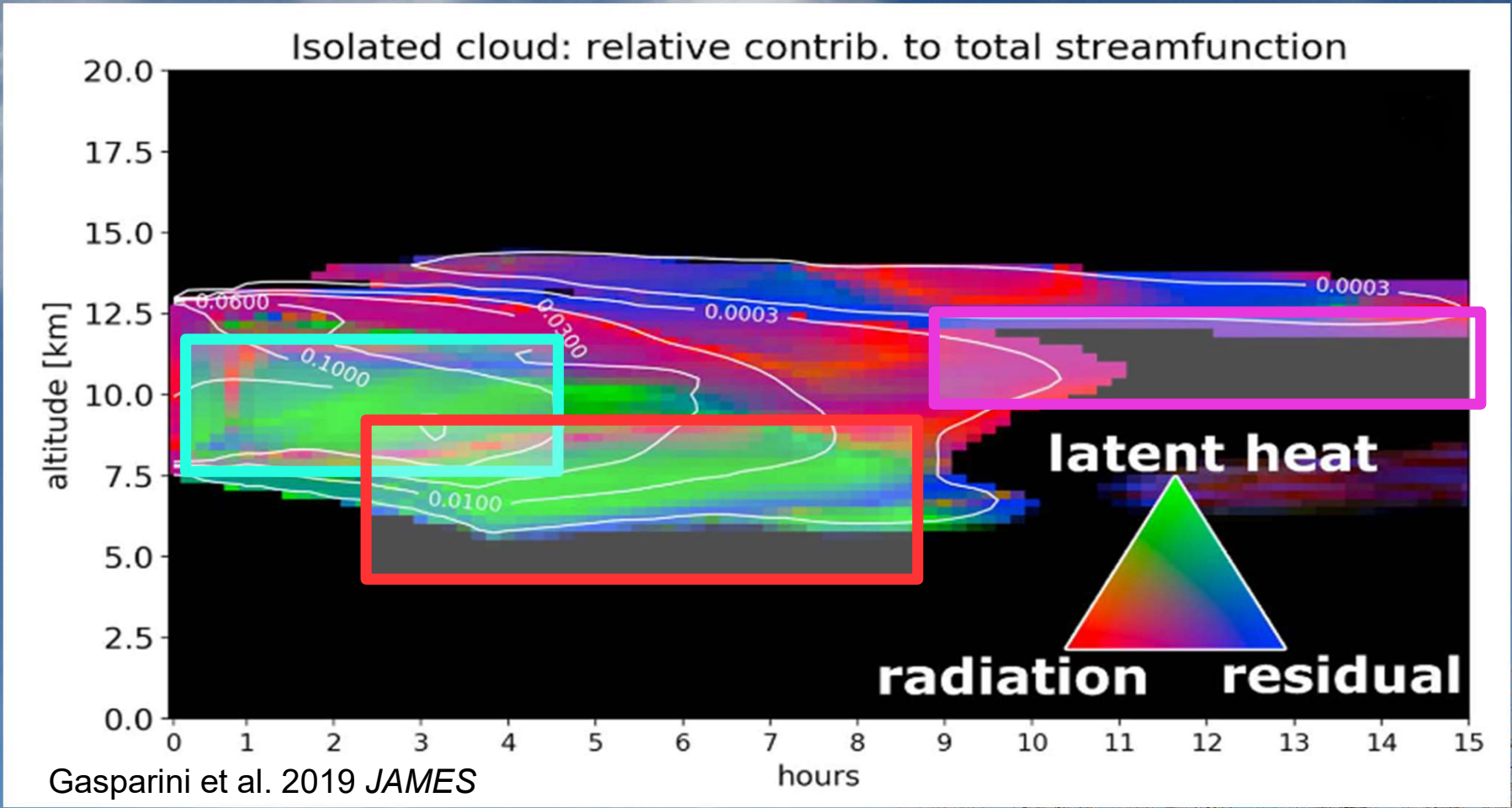
formation

development

dissipation

Radiative heating modulates and is determined by these processes.

$T,$
 ρ



formation

development

dissipation

t

PART 1

ice-nucleating
particles



vertical
velocity



ice
nucleation

PART 1

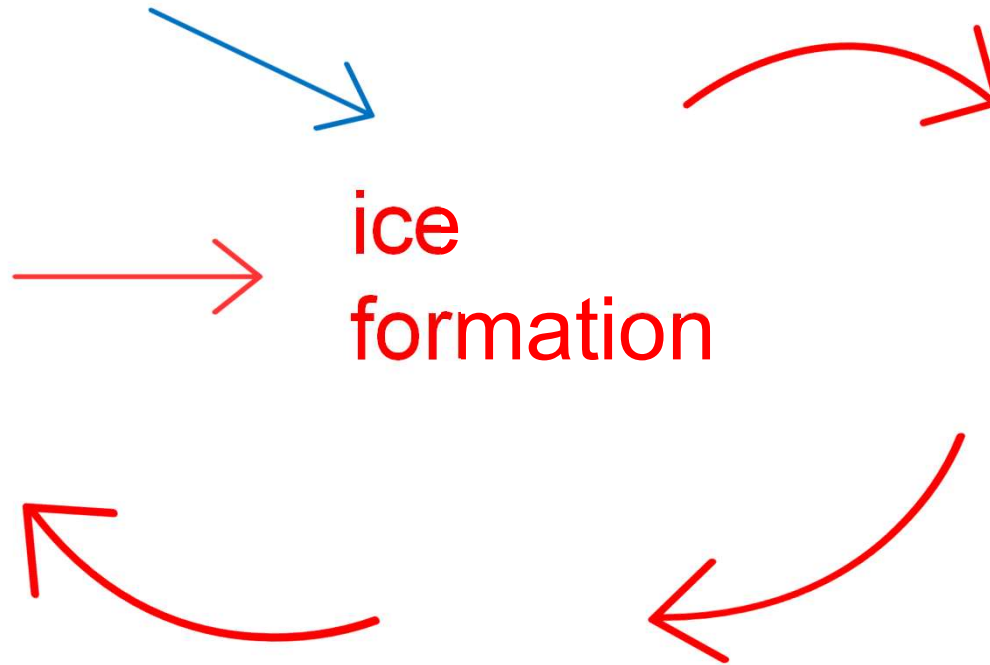
PART 2

ice-nucleating
particles

vertical
velocity

ice
formation

cloud-radiative
heating



PART 1

PART 2

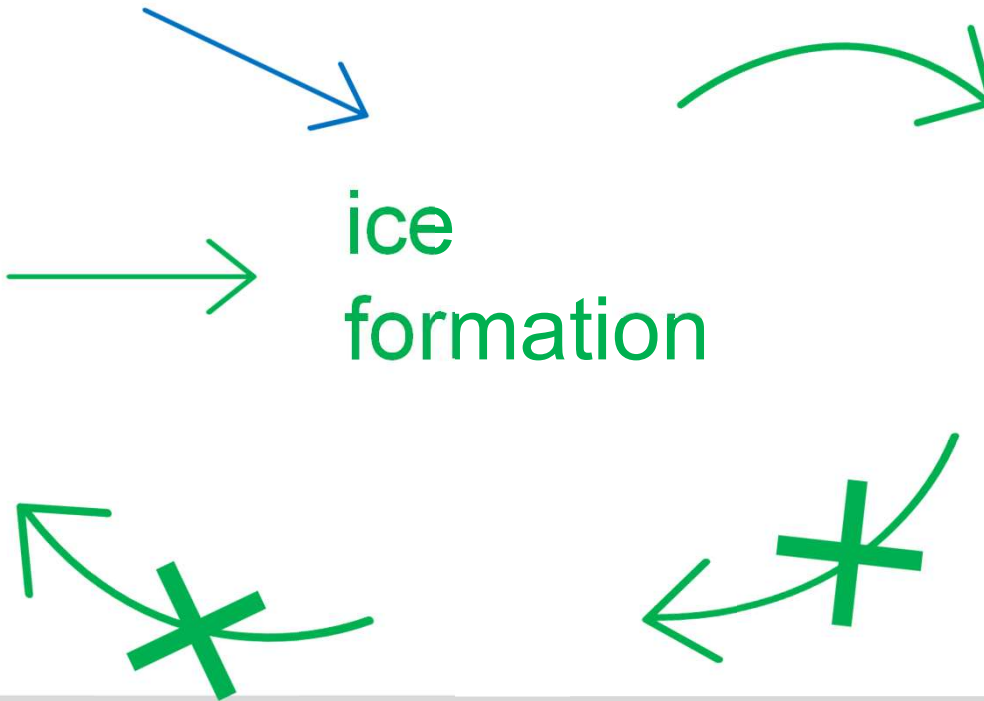
PART 3

ice-nucleating
particles

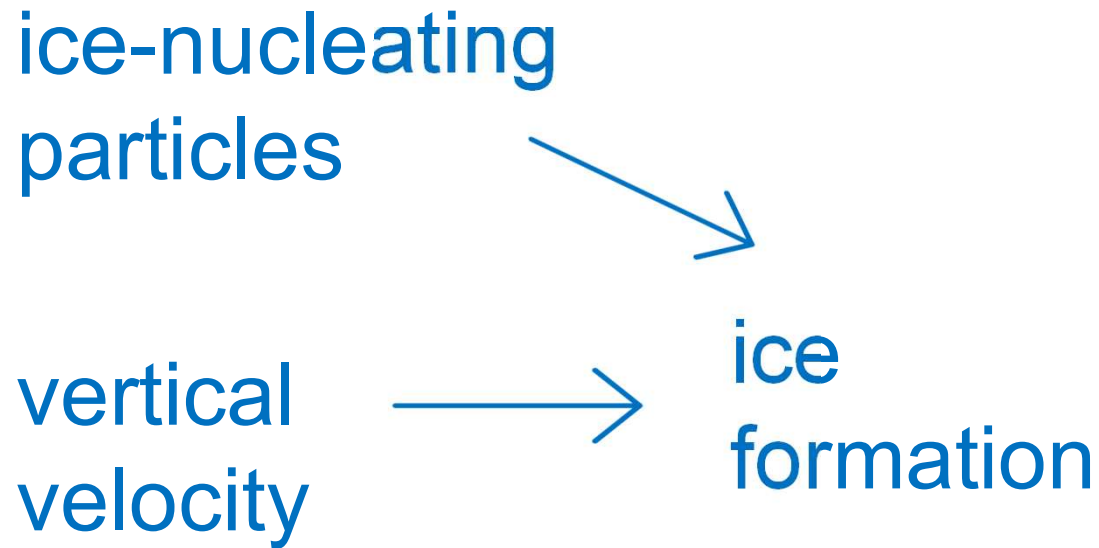
vertical
velocity

ice
formation

cloud-radiative
heating

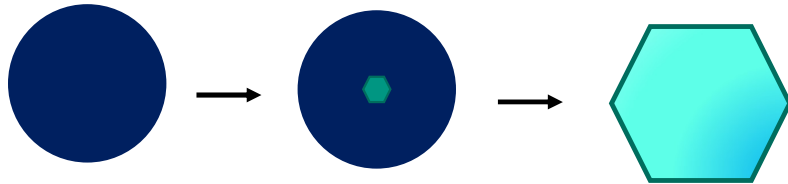


PART 1

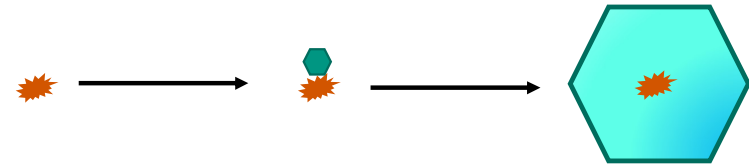


How do we represent ice formation in the atmosphere?

How do we represent ice formation in the atmosphere?



Homogeneous nucleation



Heterogeneous nucleation

Supersaturation is key

$$J(s_i) \frac{ds_i}{dt} = \text{source from expansion cooling} - \text{sink to ice crystal growth}$$

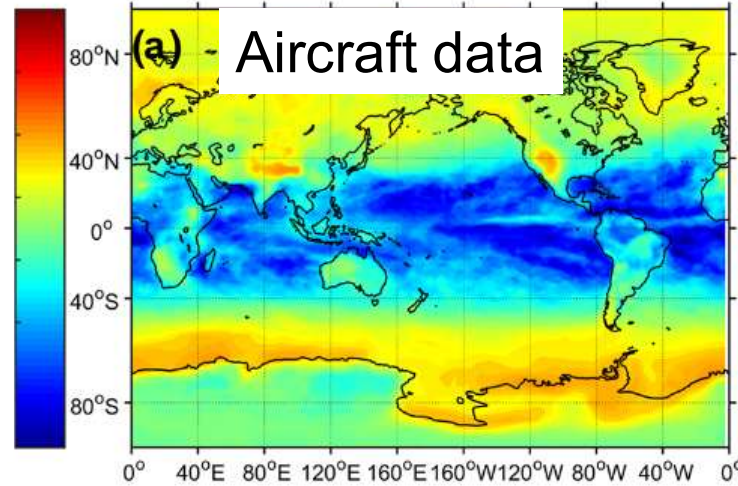
$$N_i = f(v_z, T, N_{\text{Sulf}}, N_{\text{BC}}, N_{\text{Dust}}, N_{\text{Org}})$$

Ice crystal numbers vary dramatically in formulations from different sources.

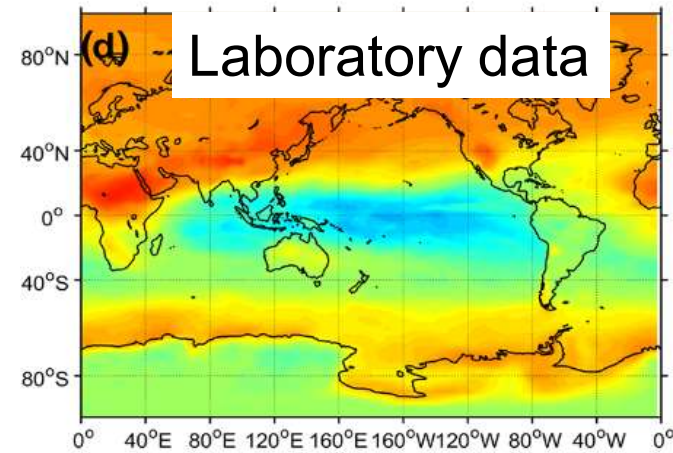
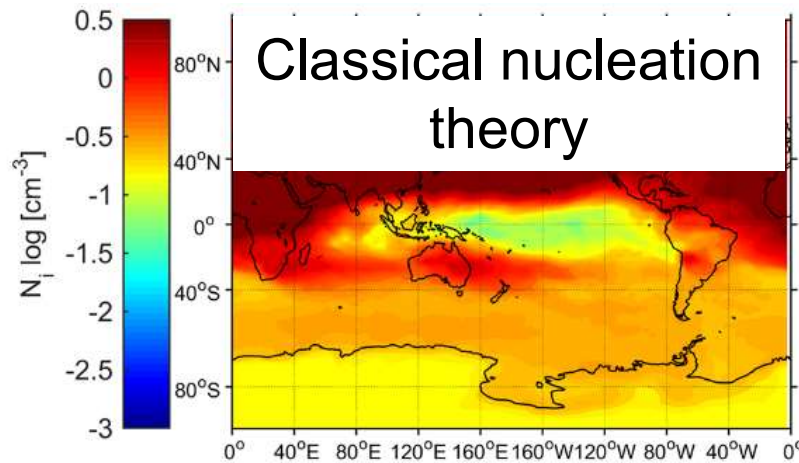
*more ice crystals
per volume air*



*less ice crystals
per volume air*



Annual average from the
Community Atmosphere Model 5.0
at the 232-hPa level



How can we better understand these ice crystal number differences
in a computationally efficient and interpretable manner?

$$N_i = f(v_z, T, N_{\text{Sulf}}, N_{\text{BC}}, N_{\text{Dust}}, N_{\text{Org}})$$

Automatic differentiation

$$\left(N_i, \frac{\partial N_i}{\partial v_z}, \frac{\partial N_i}{\partial T}, \frac{\partial N_i}{\partial N_{\text{Sulf}}}, \frac{\partial N_i}{\partial N_{\text{BC}}}, \frac{\partial N_i}{\partial N_{\text{Dust}}}, \frac{\partial N_i}{\partial N_{\text{Org}}} \right) =$$

$$f(v_z, T, N_{\text{Sulf}}, N_{\text{BC}}, N_{\text{Dust}}, N_{\text{Org}})$$

We can use sensitivities to define 'attribution metrics'. 

sensitivity

input variance

$$\xi_{x_j}^{(N_i)} = \frac{\overline{\left(\frac{\partial N_i}{\partial x_j}\right)^2} \sigma_{x_j}^2}{\sum_{j=1}^J \overline{\left(\frac{\partial N_i}{\partial x_j}\right)^2} \sigma_{x_j}^2}$$

0 % - x_j is not linked to temporal variability in N_i

100 % - all temporal variability in N_i is linked to x_j

x_j

grid cells colored by input with largest attribution

 Vertical velocity

v_z

 Sulfate particles

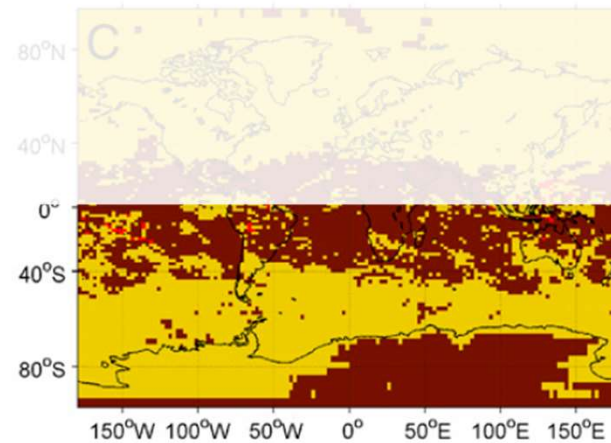
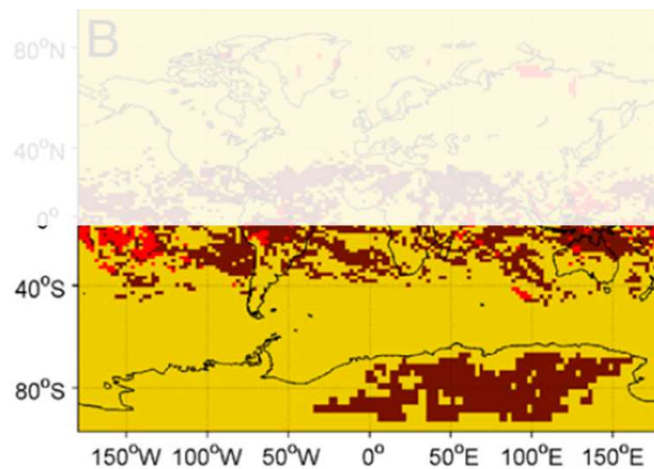
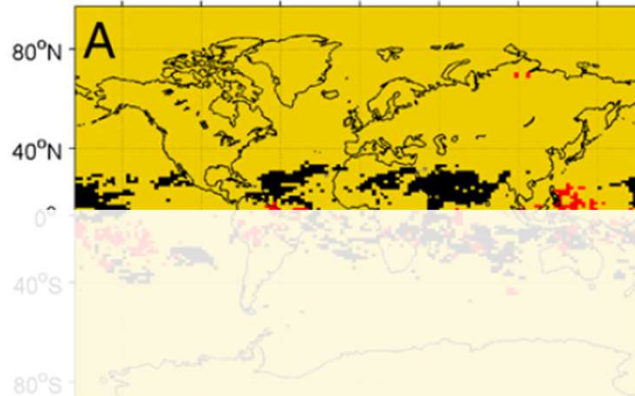
N_{Sulf}

Variability in vertical velocities is crucial to represent ice nucleation.

N_{Dust}

 Large dust

N_{Dust}



We can use sensitivities to define 'attribution metrics'.

sensitivity

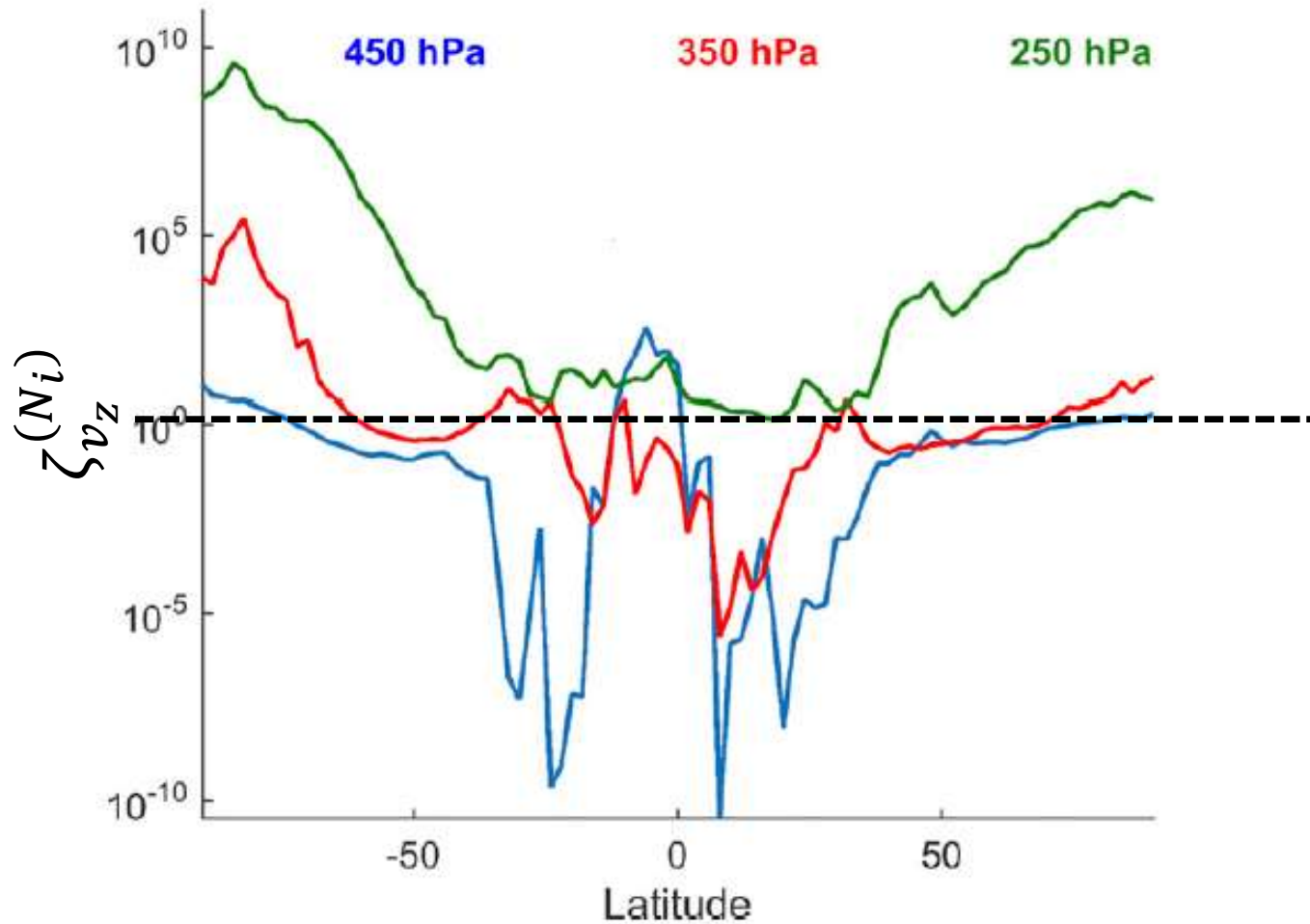
$$\zeta_{x_j}(N_i) = \frac{\overline{\left(\frac{\partial N_i}{\partial x_j}\right)^2}}{\sigma_{x_j}^2} \frac{\overline{x_j^4}}{\overline{N_i}^2}$$

input variance

$\ll 1$ – fluctuations in x_j are large

$\gg 1$ – any fluctuations in x_j are amplified

$\gg 1$ – large sensitivity to v_z

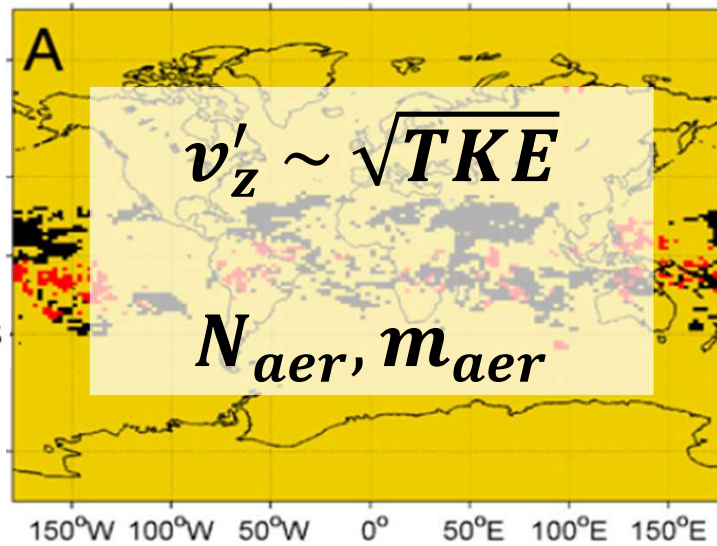


At higher latitudes and altitudes, sensitivity to vertical velocity causes its large attribution.

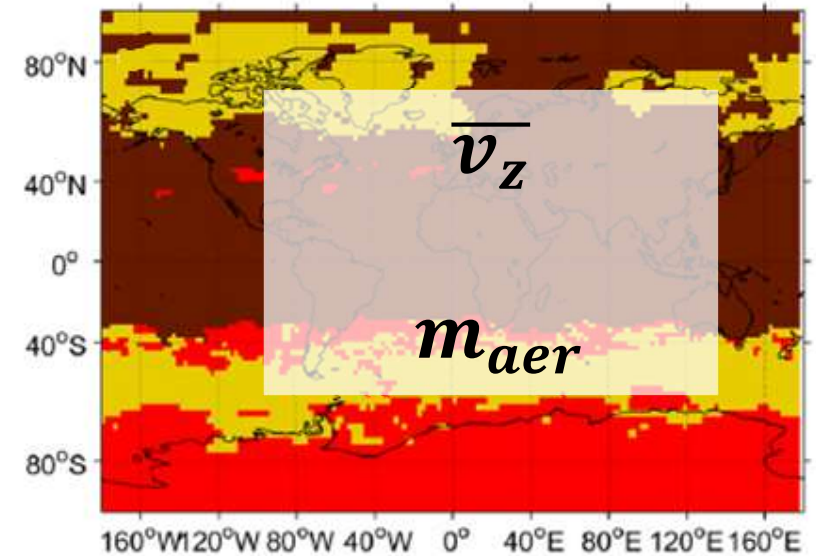
At lower latitudes and altitudes, variance in vertical velocity cause its large attribution.

Subgrid-scale variability in v_z and aerosol module alter model attributions.

- Vertical velocity
 v_z
- Sulfate particles
 N_{Sulf}
- Small dust particles
 N_{Dust}
- Large dust
 N_{Dust}



Community Atmosphere Model 5.0



Goddard Earth-Observing Model

PART 1

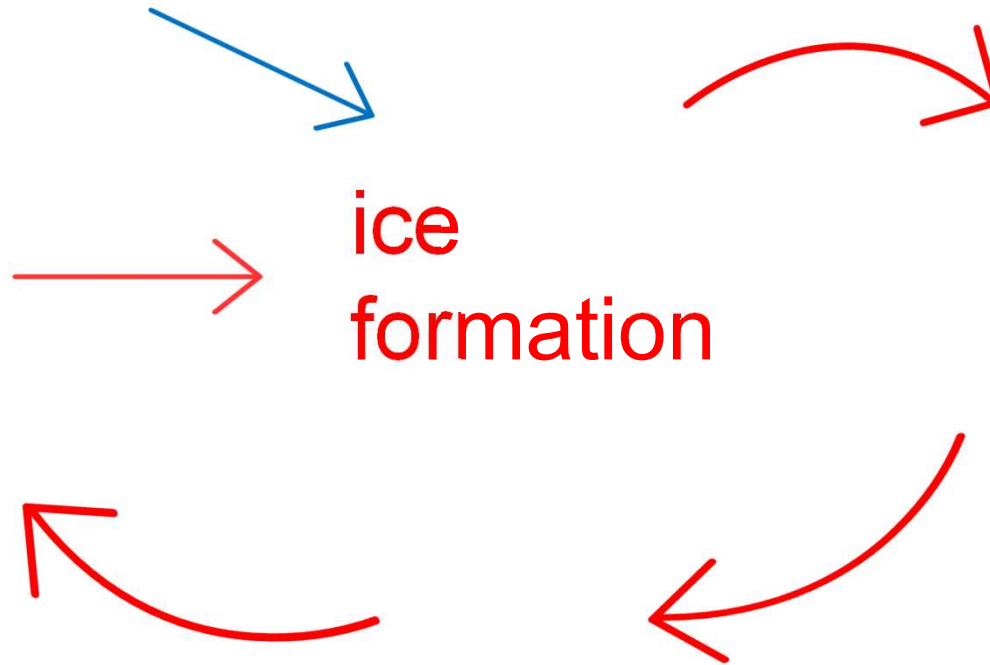
PART 2

ice-nucleating
particles

vertical
velocity

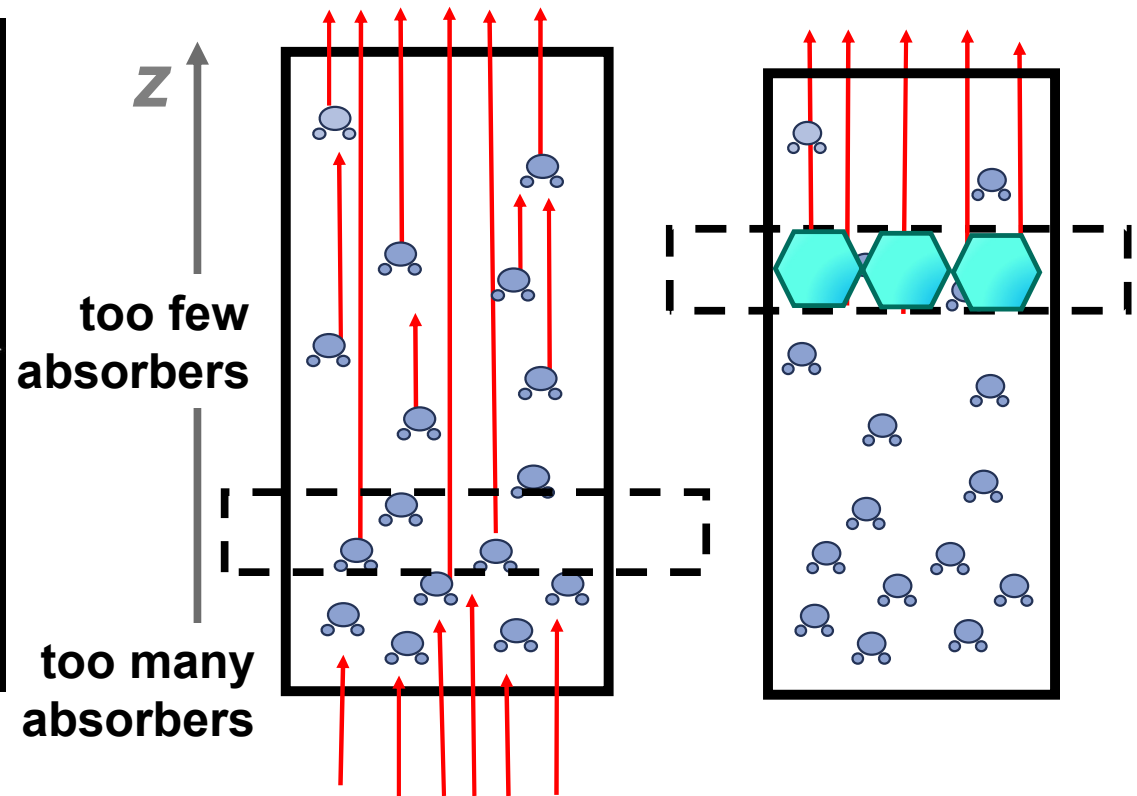
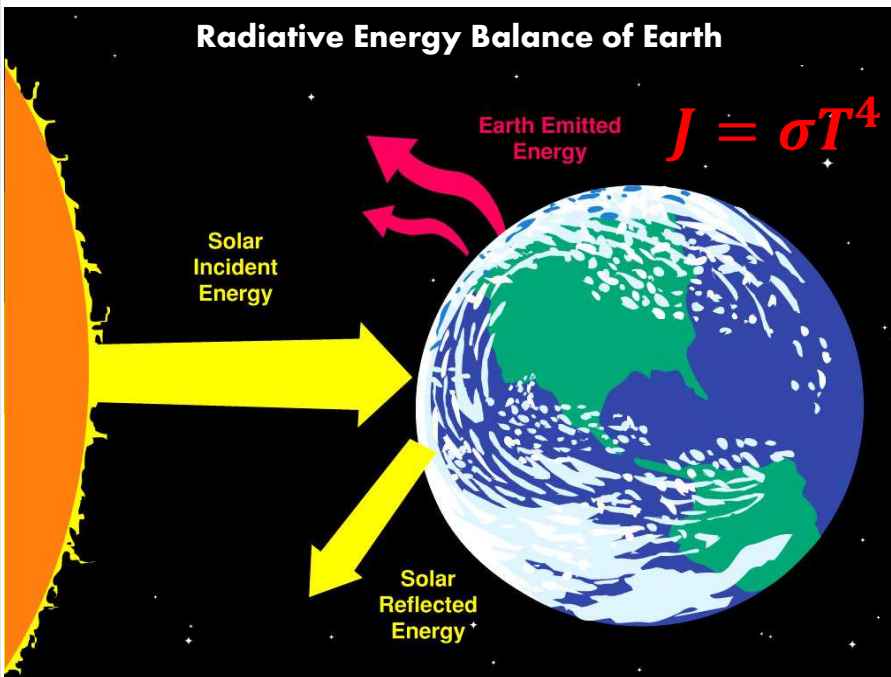
ice
formation

cloud-radiative
heating



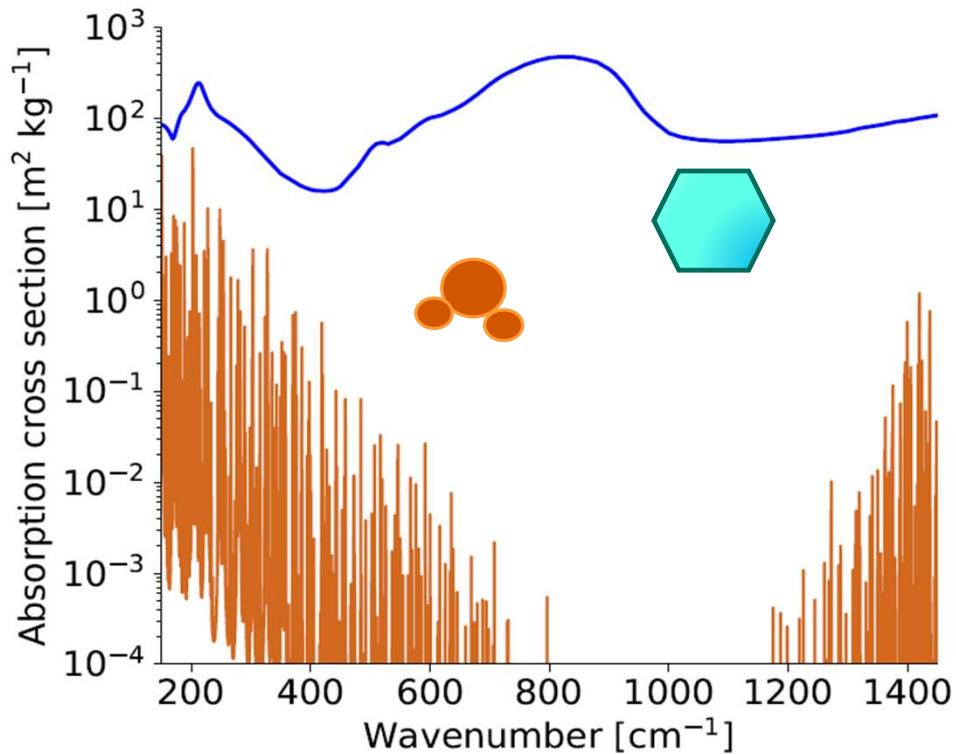
Ice cloud emits less radiation than clear sky.

less emitted infrared radiation

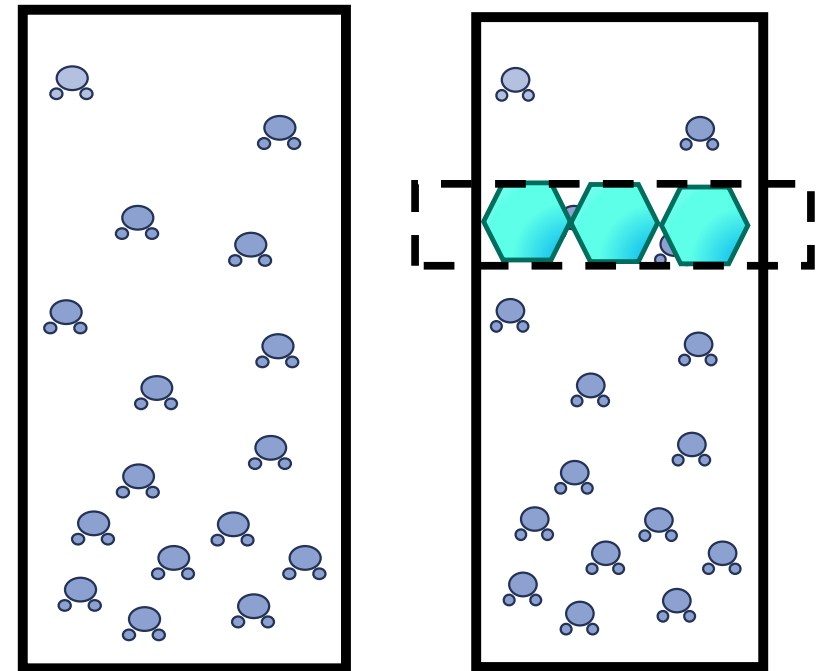


Ice cloud absorbs more radiation than clear sky.

less emitted infrared radiation
and
strong atmospheric heating



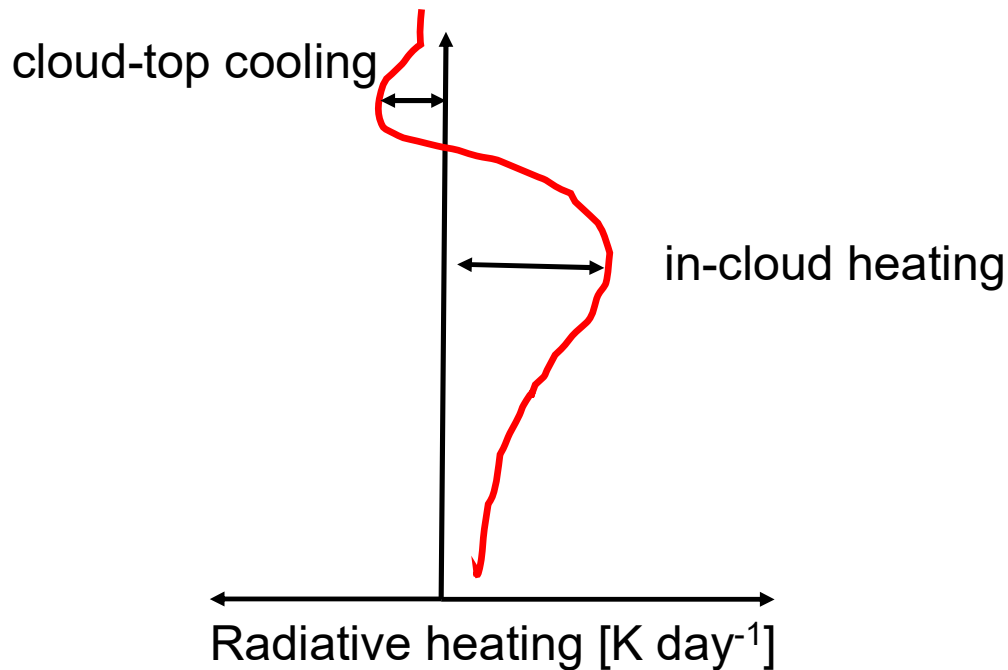
Ice data from Warren and Brandt 2008; wv data from HITRAN.



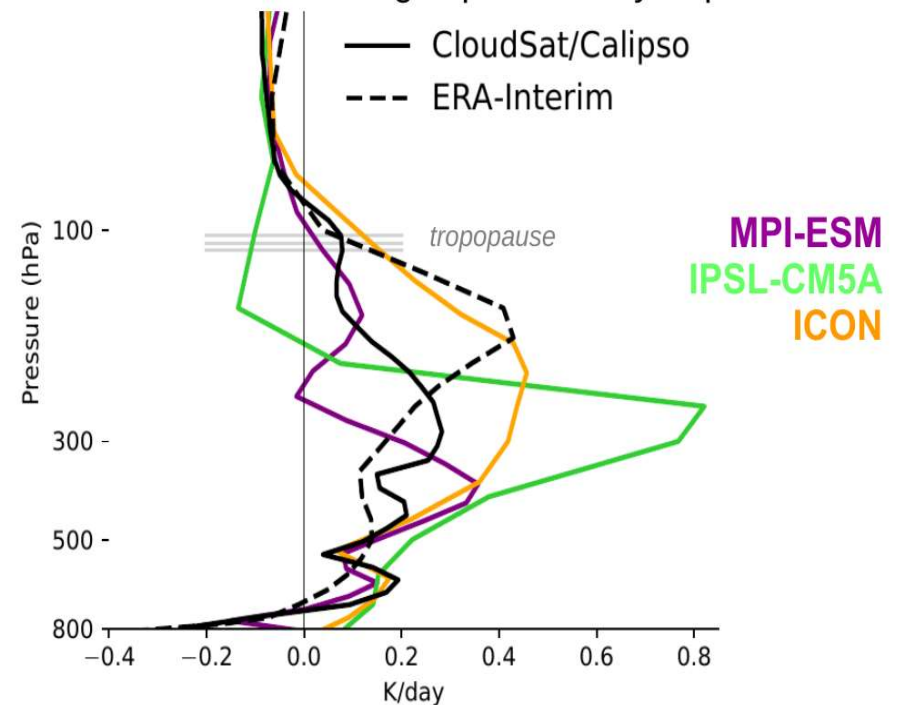
less emitted infrared radiation

and

strong atmospheric heating



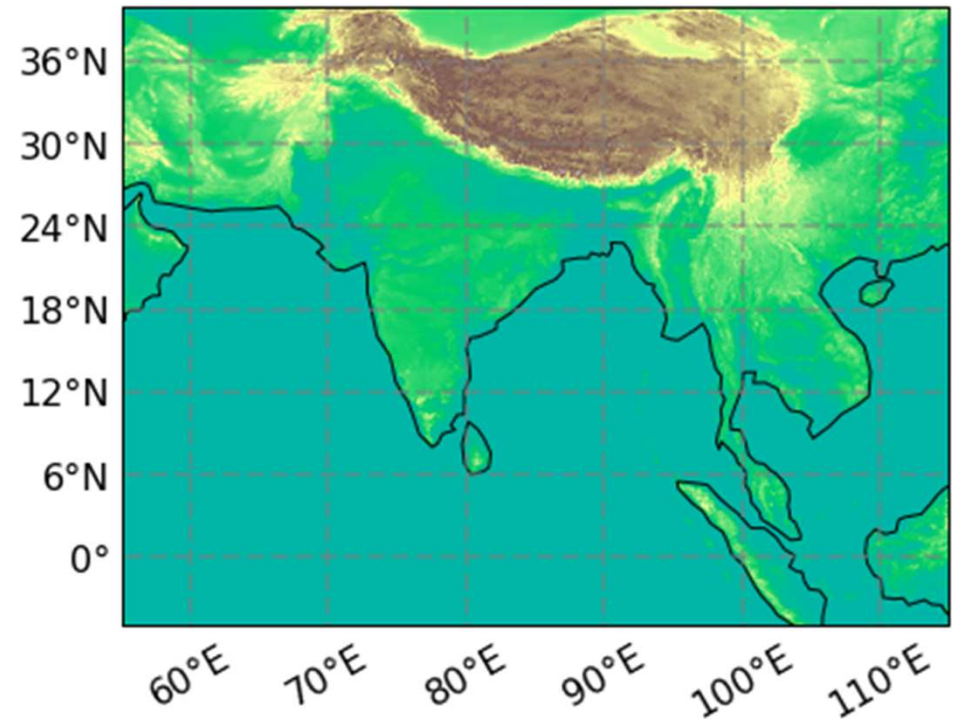
Cloud-radiative heating in present-day tropics



A. Voigt, N. Albern, G. Papavasileiou (2019) *J. Clim.*

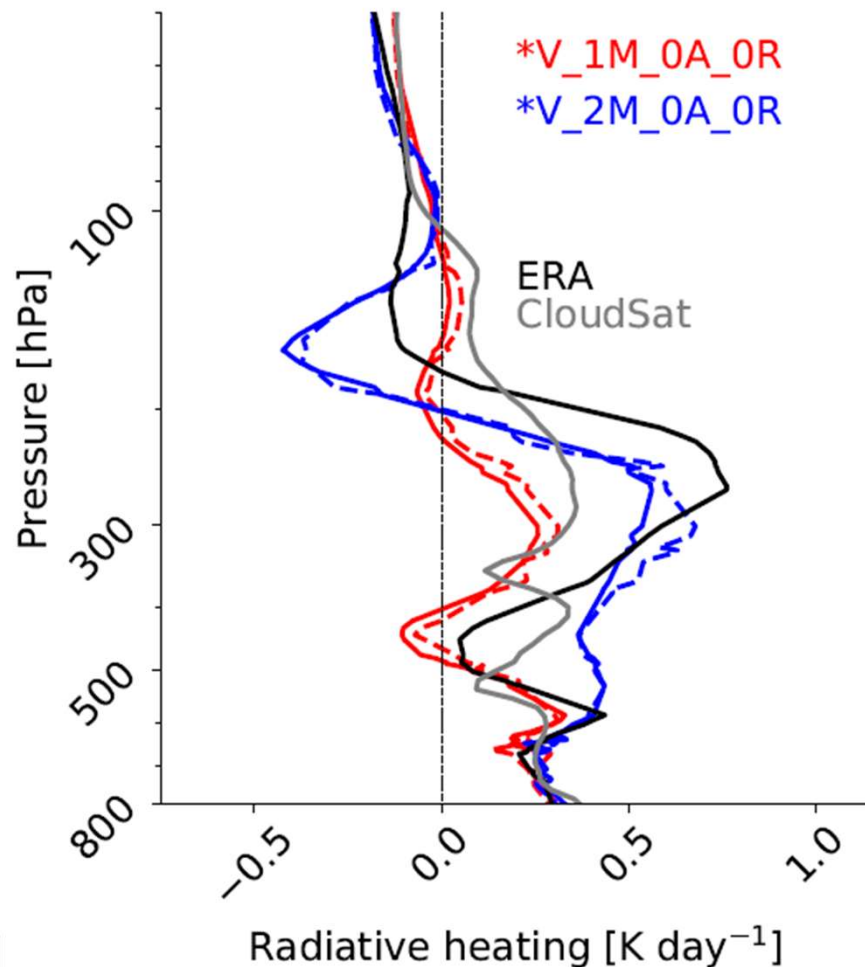
We investigate four microphysical switches in storm-resolving simulations.

- M** One-moment or two-moment
- V** Default or higher vertical resolution
- A** Aerosol dependence or not
- R** Consistent size (effective Radius) of ice crystals between microphysics and radiation or not



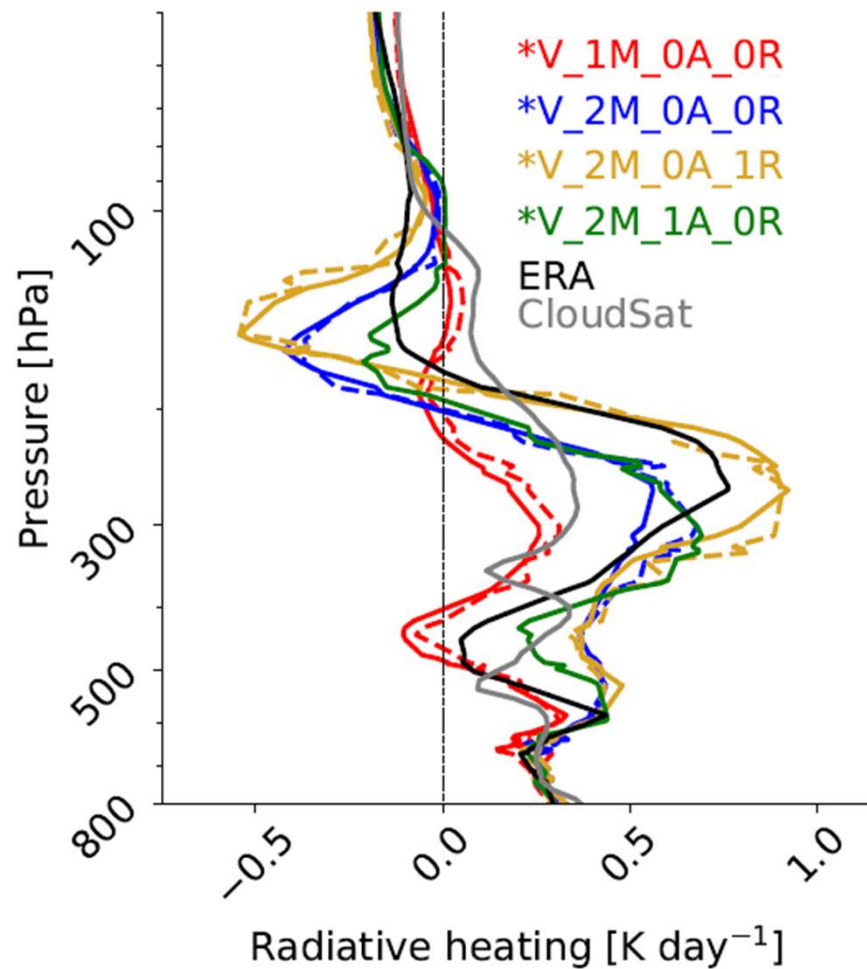
Icosahedral Nonhydrostatic Model, 2.5-km equivalent resolution, 3 days of simulation, 24-second time step

M – microphysics **V** – vert res **R** – ice crystal size **A** - aerosol

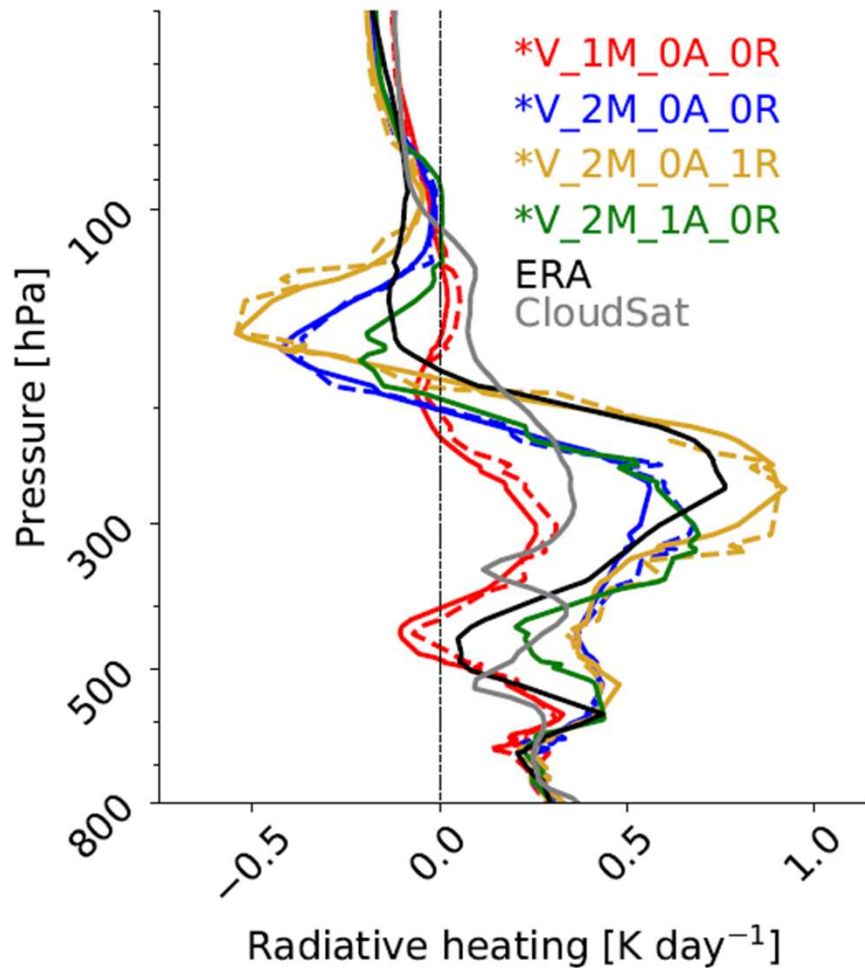


- Reanalysis and satellite profiles do not agree.
- **Vertical resolution has little impact.**
- **In-cloud heating is almost 2 times larger from 1M → 2M.**
- **Cloud-top cooling is almost 10 times larger.**

M – microphysics **V** – vert res **R** – ice crystal size **A** - aerosol



- Heating-cooling dipole increases by a factor of 2 with consistent ice crystal size.
- Heating-cooling dipole decreases above 200 hPa and below 400 hPa with aerosol dependence.



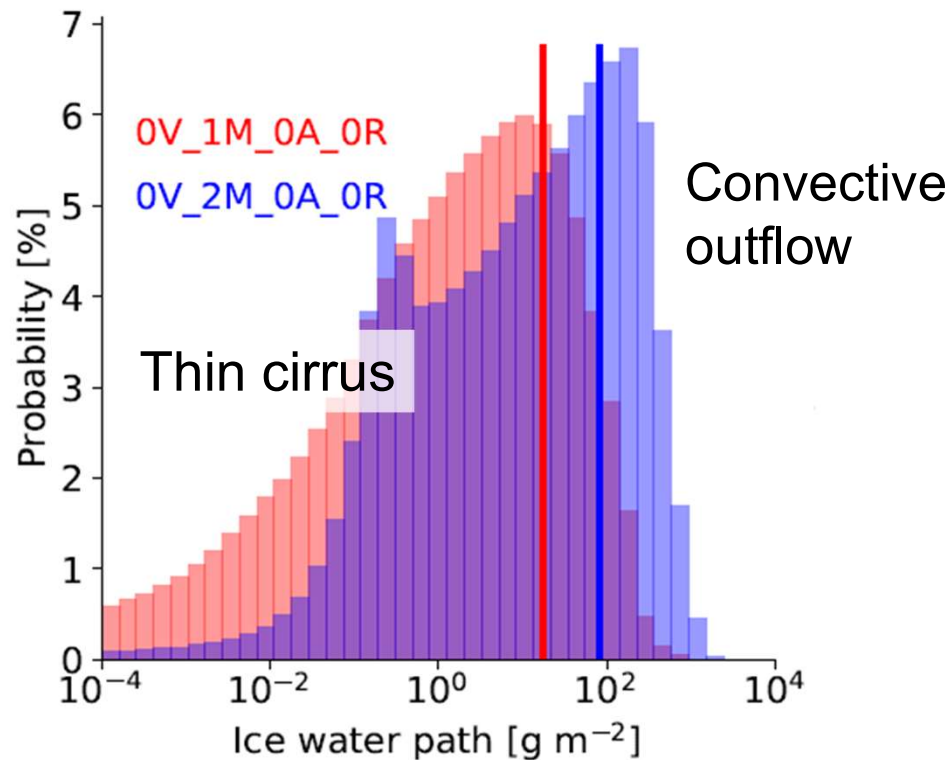
- **We can kind of generate *whatever* we want with ice microphysics switches...**

- Cloud-top cooling is almost 10 times larger.

- Heating-cooling dipole increases by a factor of 2 with consistent ice crystal size.

- Heating-cooling dipole **de ... but we can also understand why.**
an aerosol dependence.

In-cloud heating is almost 2 times larger from 1M → 2M.



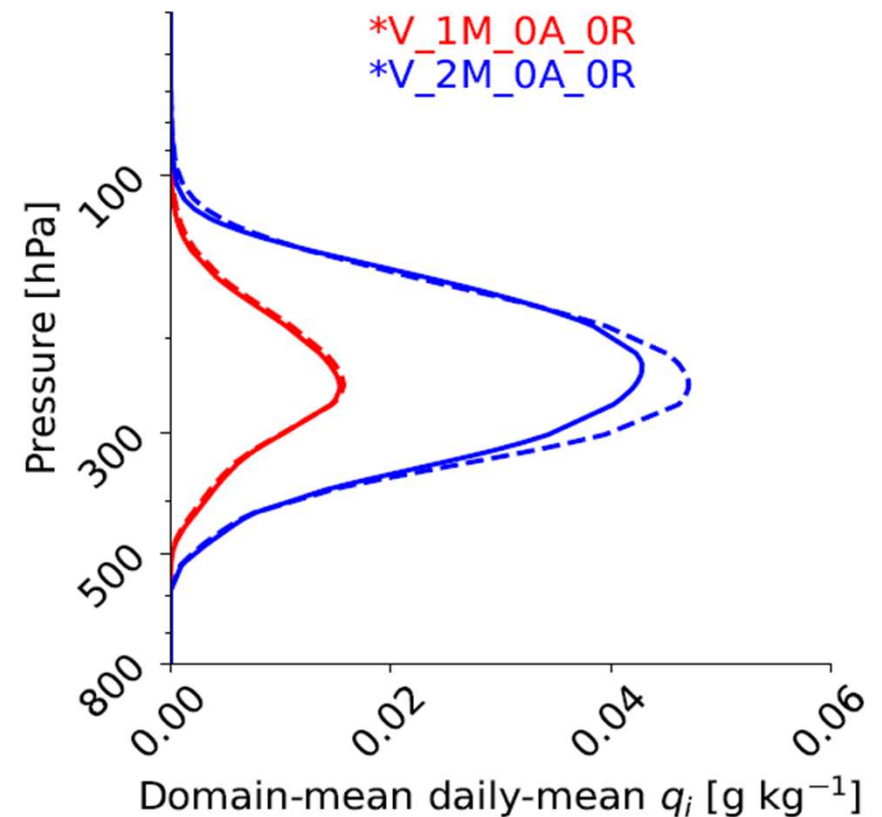
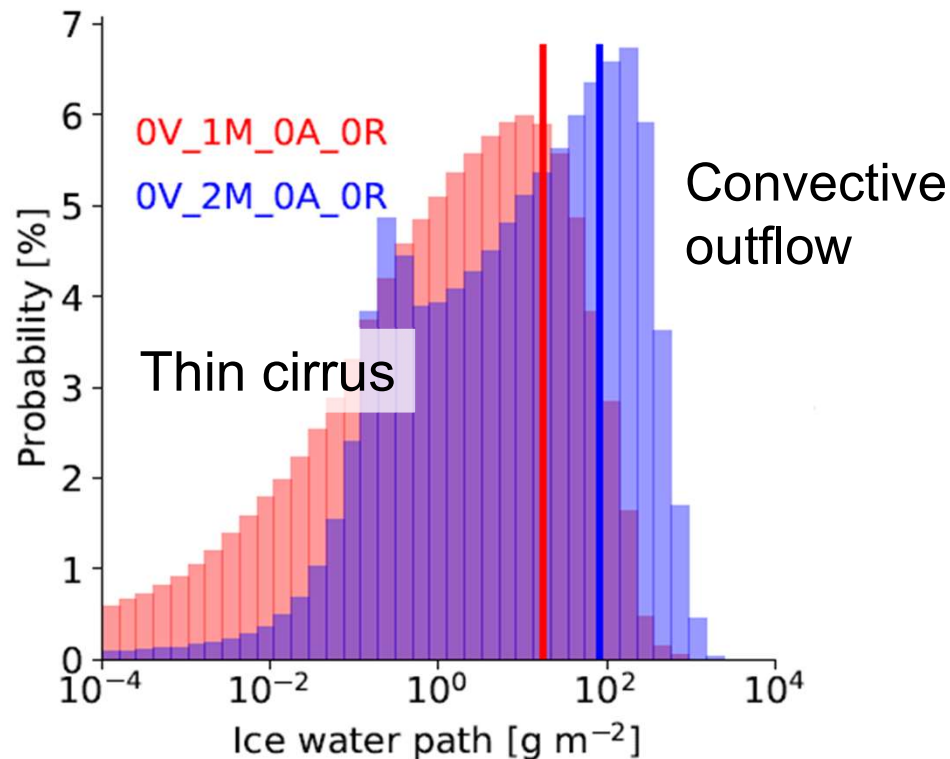
$$\frac{H_{2M}}{H_{1M}} \sim \frac{1 - \exp[-\kappa_{i,2M} \int \rho_{i,2M} dz']}{1 - \exp[-\kappa_{i,1M} \int \rho_{i,1M} dz']}$$

$$\sim \frac{1 - \exp[-100 \text{ m}^2 \text{ kg}^{-1} \cdot 0.08 \text{ kg m}^{-2}]}{1 - \exp[-100 \text{ m}^2 \text{ kg}^{-1} \cdot 0.008 \text{ kg m}^{-2}]}$$

$$\sim 2$$

Unimodal versus bimodal IWP distribution

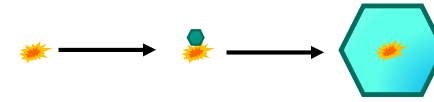
In-cloud heating is almost 2 times larger from 1M → 2M.



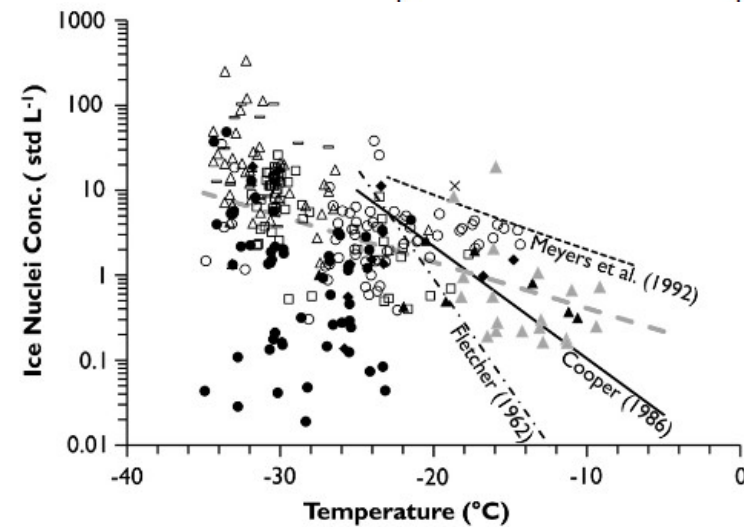
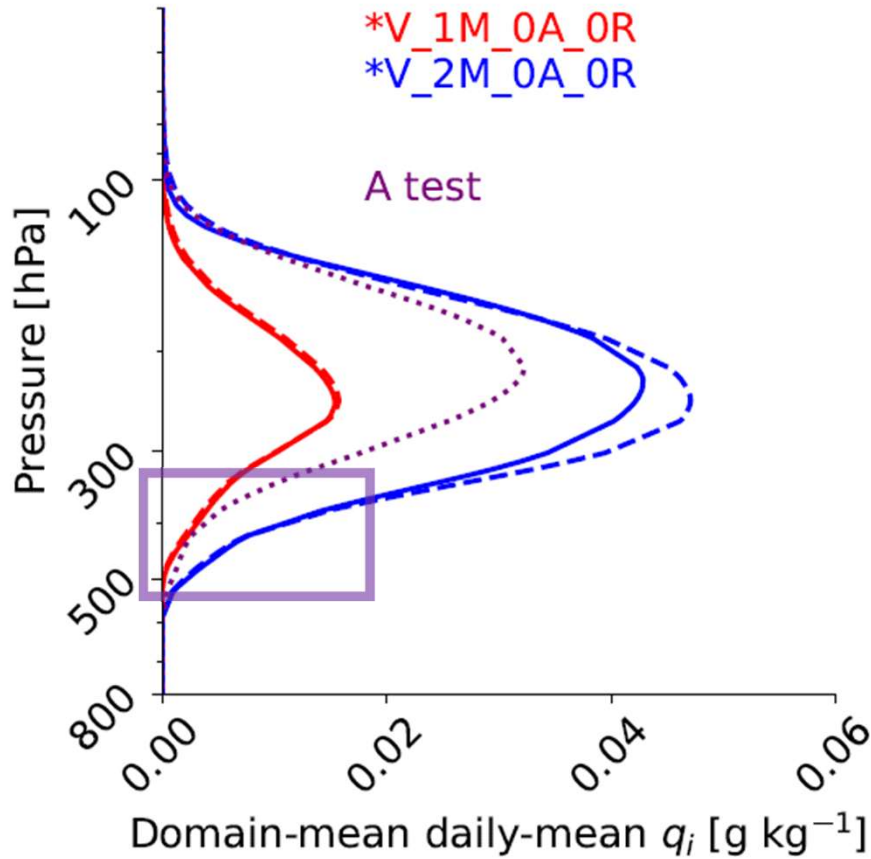
Unimodal versus bimodal IWP distribution

2M generates >4 times more ice *and* over a greater depth.

Why does 2M generate >4 times more ice? *low*



$$C_{INP} = A \exp \left[- B(T - T_{min})^C \right]$$



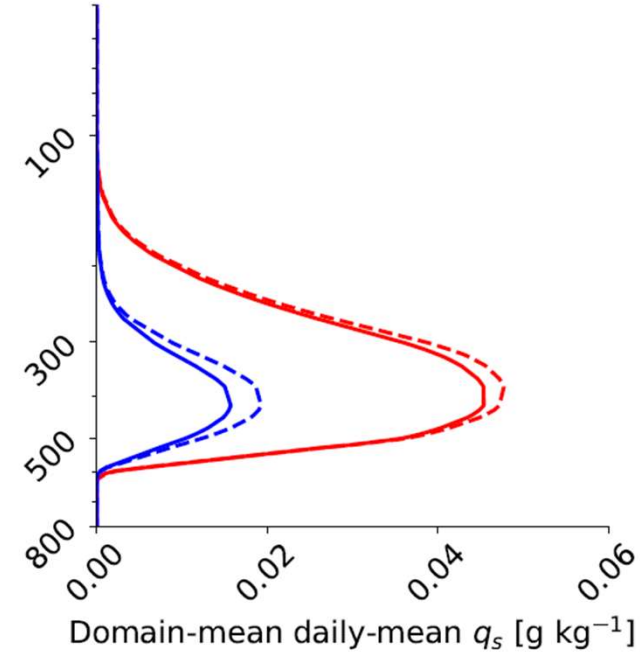
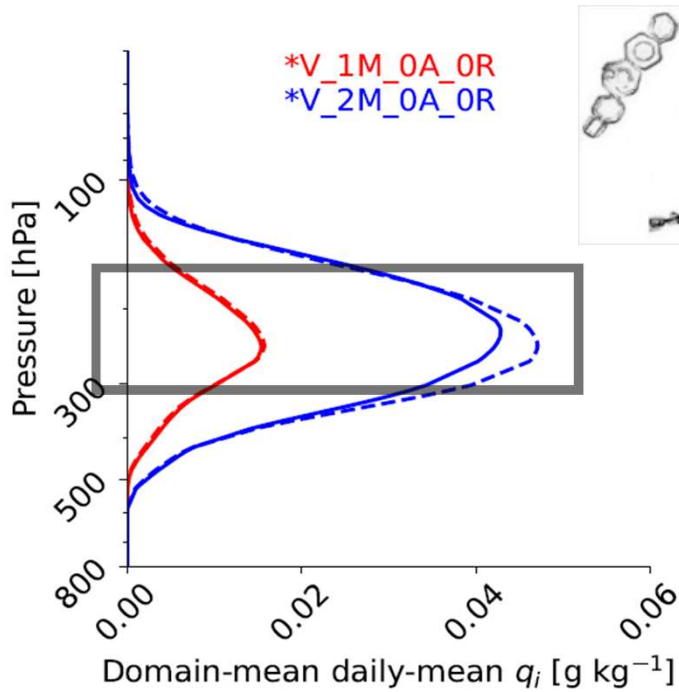
$T_{min} = 237 \text{ K}$
 $A = 2.969 \times 10^4 \text{ m}^{-3}$

$T_{min} = 237 \text{ K}$
 $A = 1.0 \times 10^2 \text{ m}^{-3}$

Parameters from Hande et al. 2015 ACP & Doms et al. 2005

Fig from DeMott et al. 2010 PNAS

Why does 2M generate >4 times more ice? *middle*



ice-to-snow conversion processes:

i + i = s (aggregation, ice autoconversion)

s + c = i (riming)

i + s = s (aggregation)

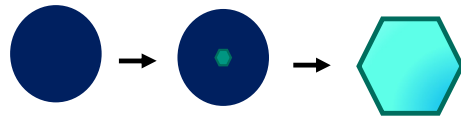
$$\frac{\partial m_a}{\partial t} \propto e_{ab} N_a m_b G_1(\delta_a, \delta_b, \theta_a, \theta_b) \quad S_{au} = C(q_i - q_{i,0})$$

$$\frac{\partial m_b}{\partial t} = -\frac{\partial m_a}{\partial t}$$

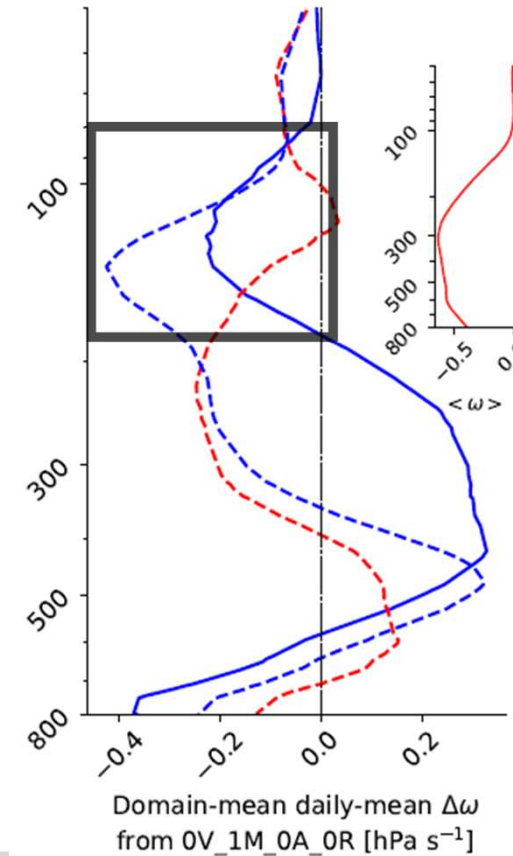
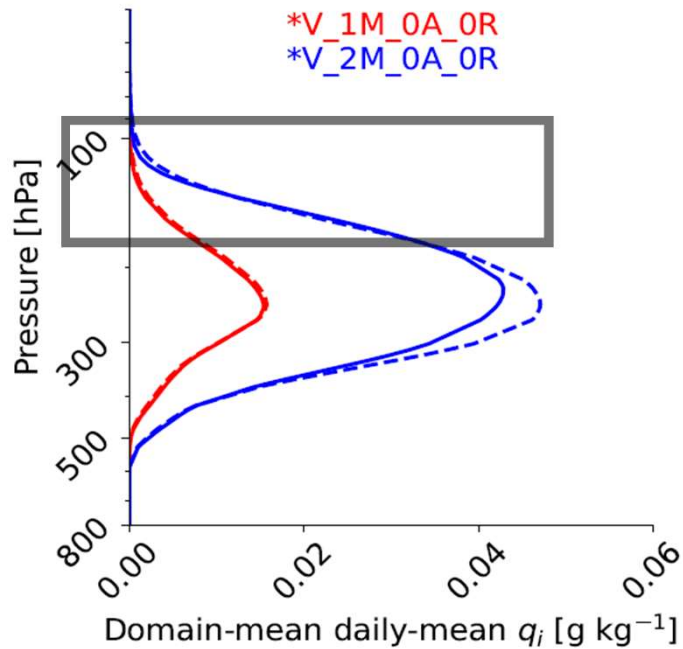
$$\frac{\partial N_a}{\partial t} \propto -e_{ab} N_a N_b G_2(\delta_a, \delta_b, \theta_a, \theta_b)$$

$$C = 10^3 \text{ s}^{-1}$$

Why does 2M generate >4 times more ice? *high*

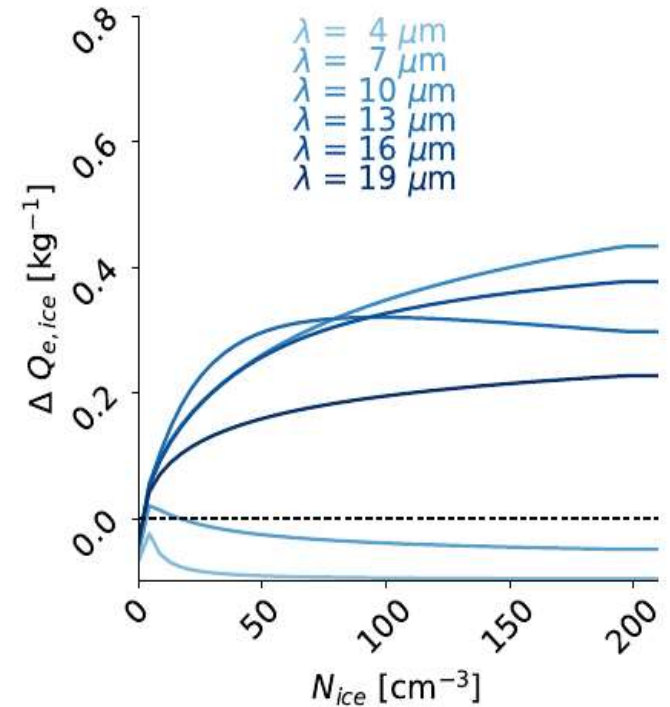
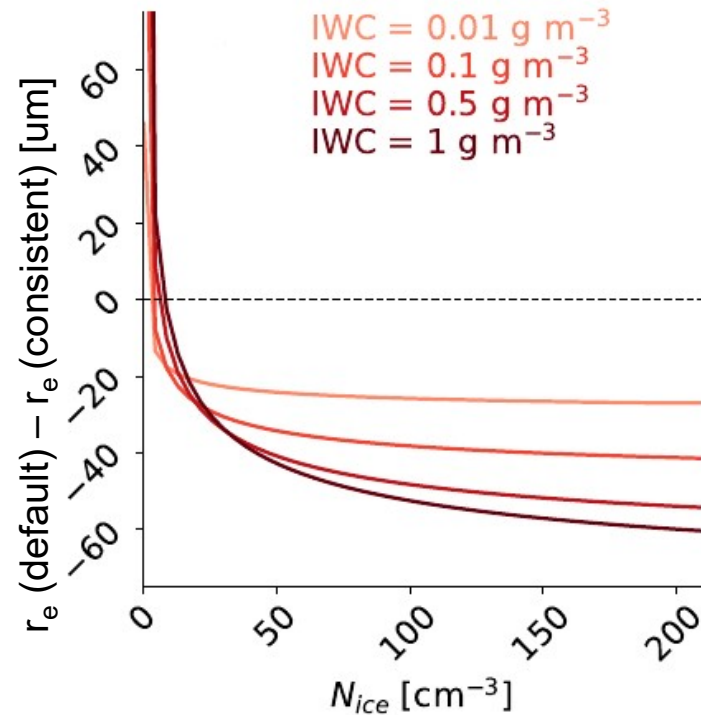
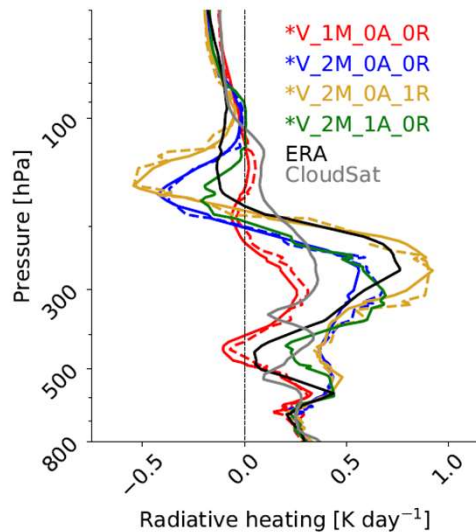


2 times larger UT mean vertical velocity in 2M



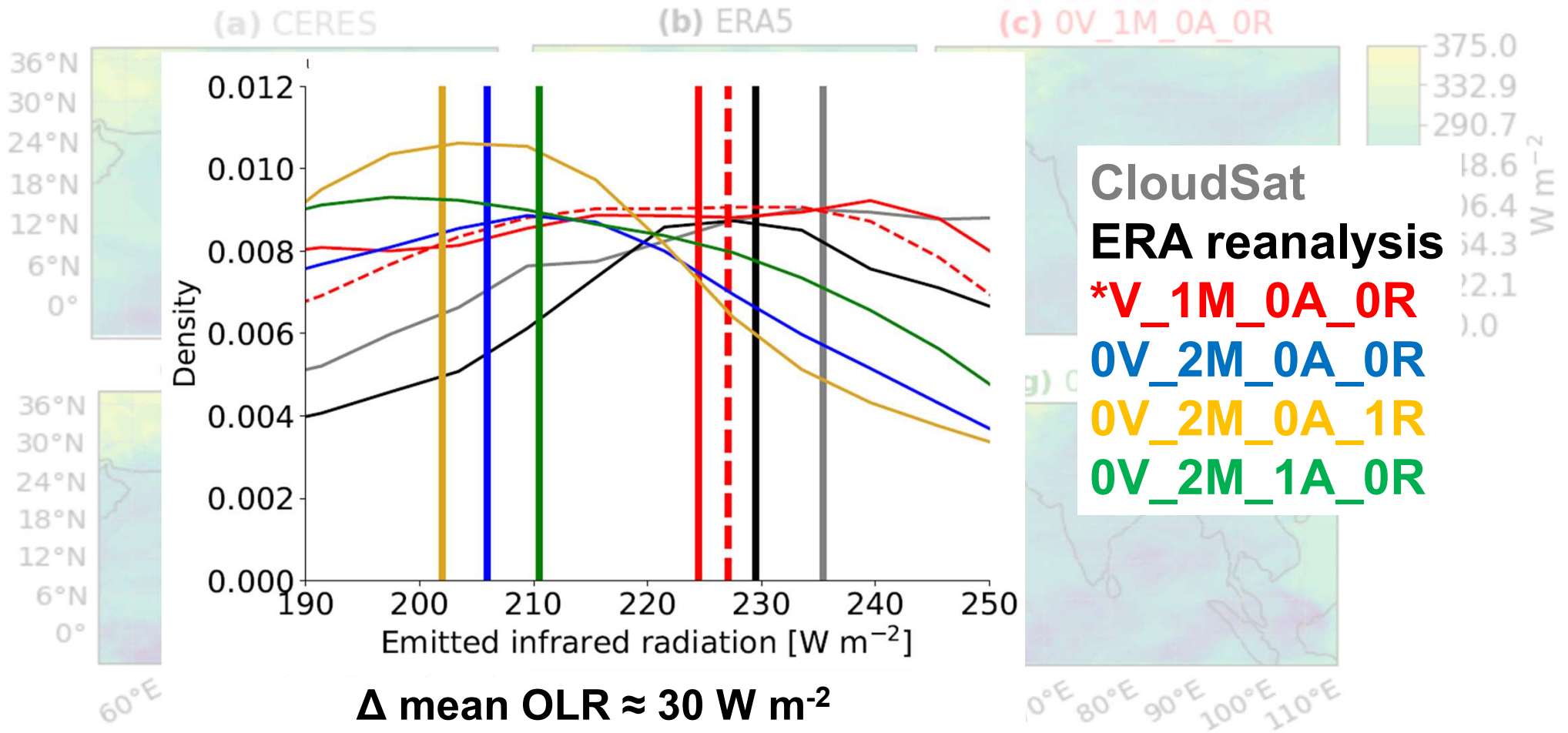
And what about the ice crystal effective radius?

$$\kappa_{i,1R} \sim Q_i \quad H_{1R} \sim 1 - \exp\left[-\kappa_{i,1R} \int \rho_{i,1R} dz'\right]$$



Here, increased extinction efficiency in the infrared window is key.

Fields of outgoing longwave radiation also change dramatically.



PART 1

PART 2

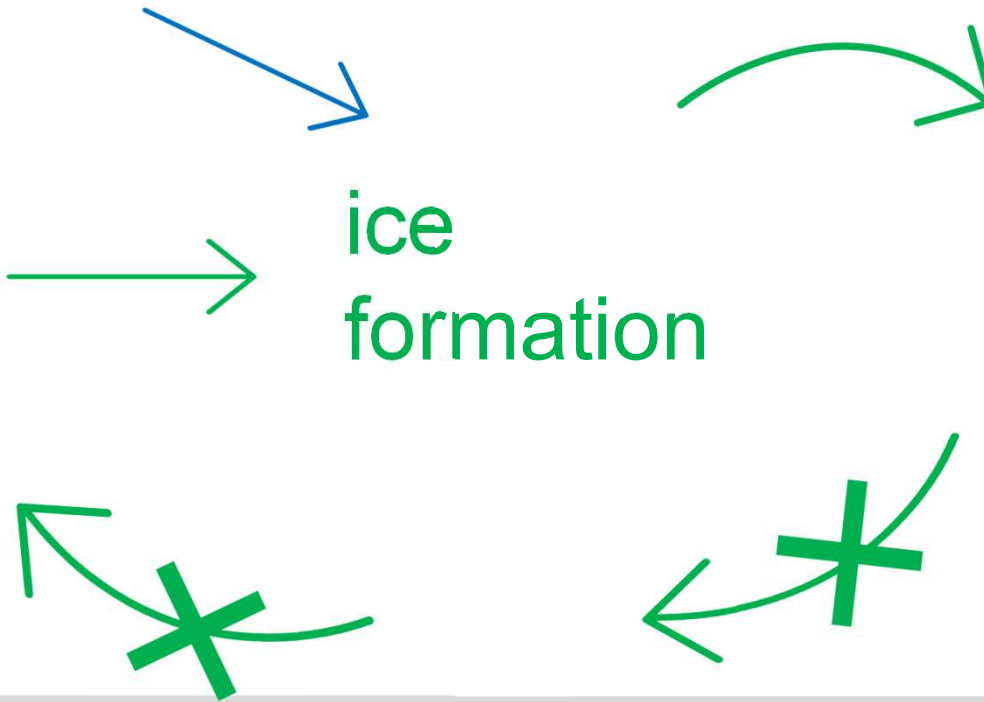
PART 3

ice-nucleating
particles

vertical
velocity

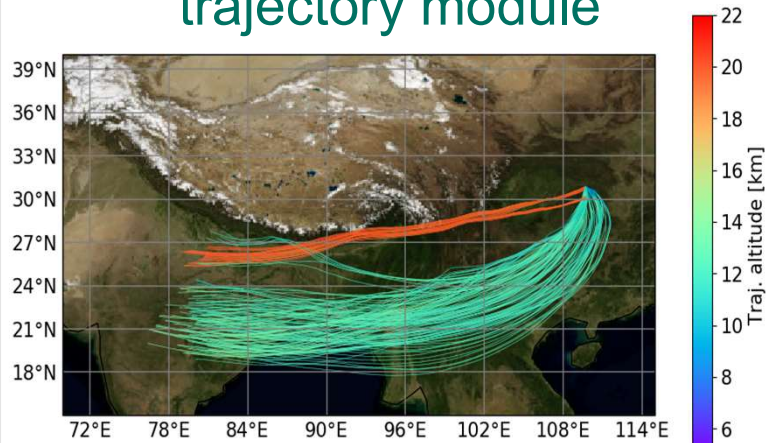
ice
formation

cloud-radiative
heating



Dynamics is fixed along Lagrangian trajectories.

LAGRANTO-based trajectory module



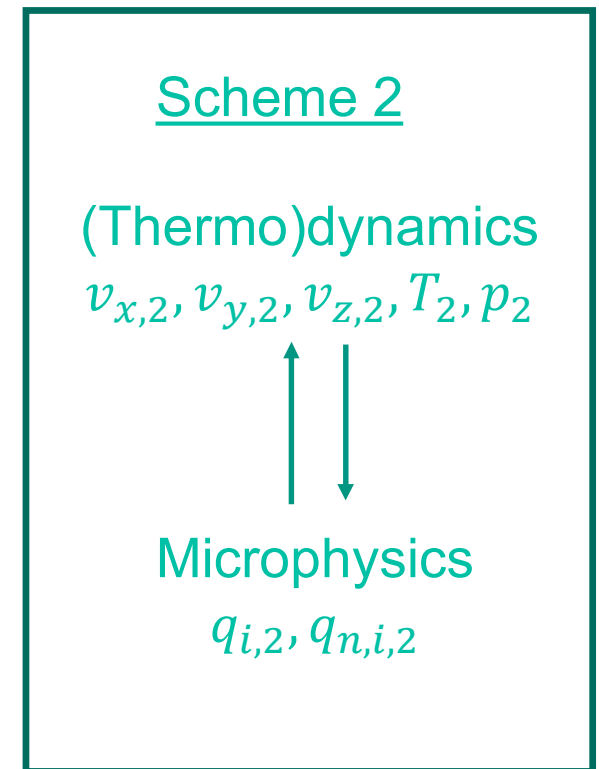
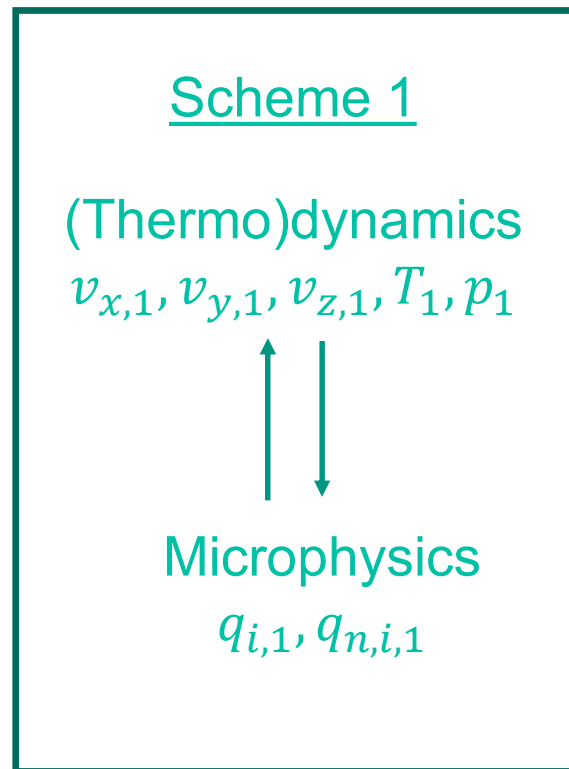
Offline radiative transfer code

CLaMS-ICE
box model

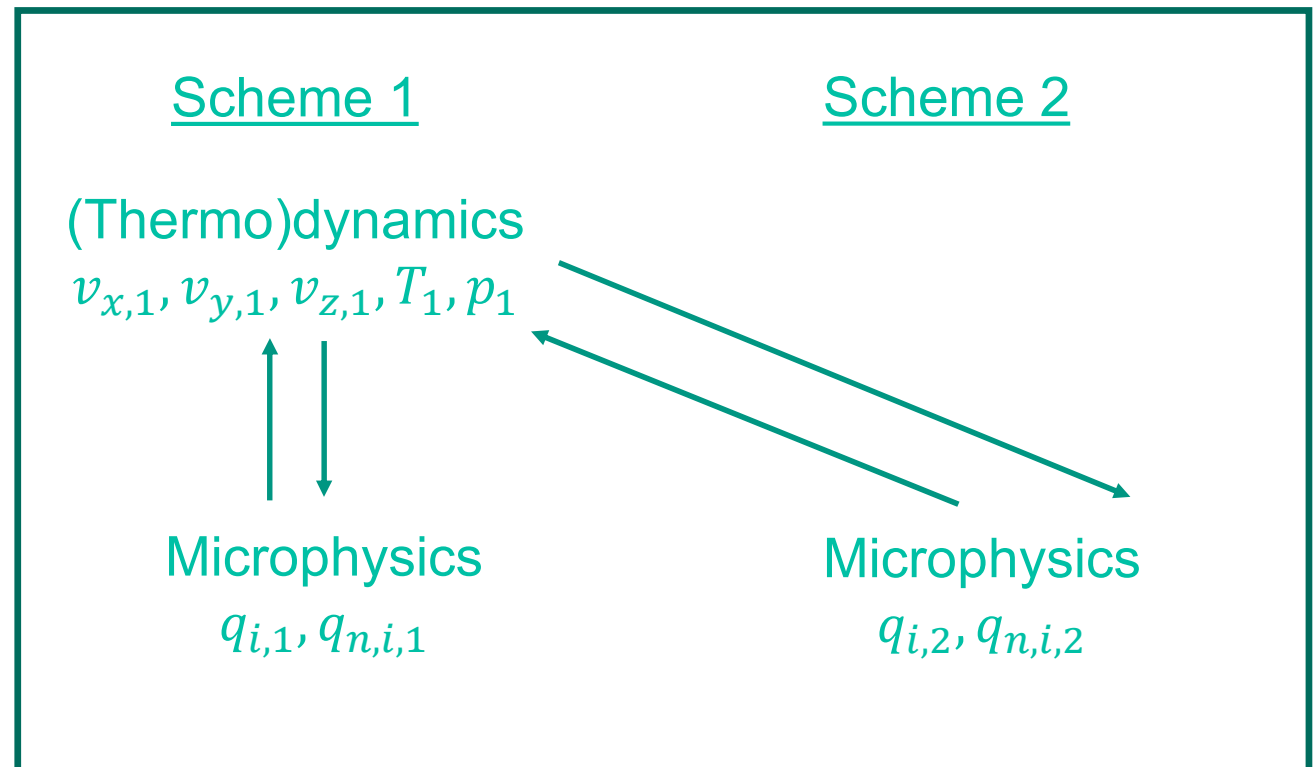


```
CALL rrtm_interface(                                     &  
    & current_date ,horizontal mesh , vertical levels ,&  
    & land fraction ,solar zenith angle ,                ,&  
    & diffuse and direct visible and near-IR albedos     ,&  
    & land emissivity ,full and half pressure levels    ,&  
    & full and half temperature levels ,vapor mixing ratio ,&  
    & liquid mixing ratio ,ice mixing ratio              ,&  
    & cloud droplet number concentration ,ice crystal number ,&  
    & cloud fraction ,trace gas concentrations           ,&  
    & aerosol concentrations                             ,&  
    & longwave and shortwave net and clear-sky fluxes  )
```

Dynamics is fixed along Lagrangian trajectories
in a variant of microphysical piggybacking.

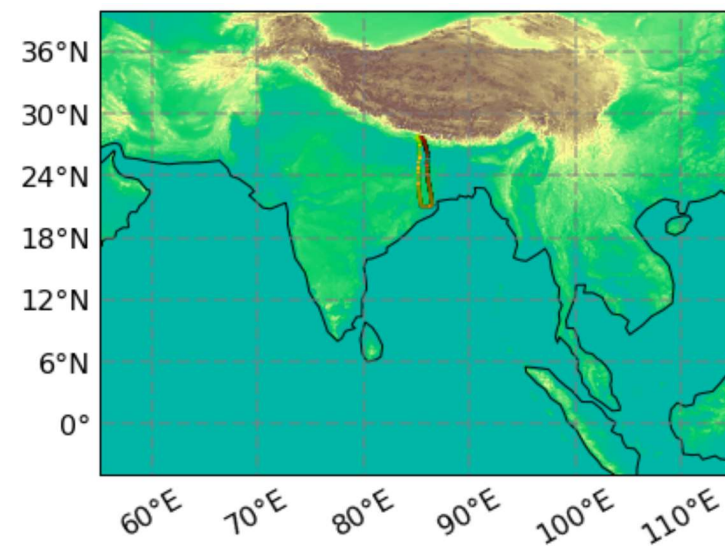
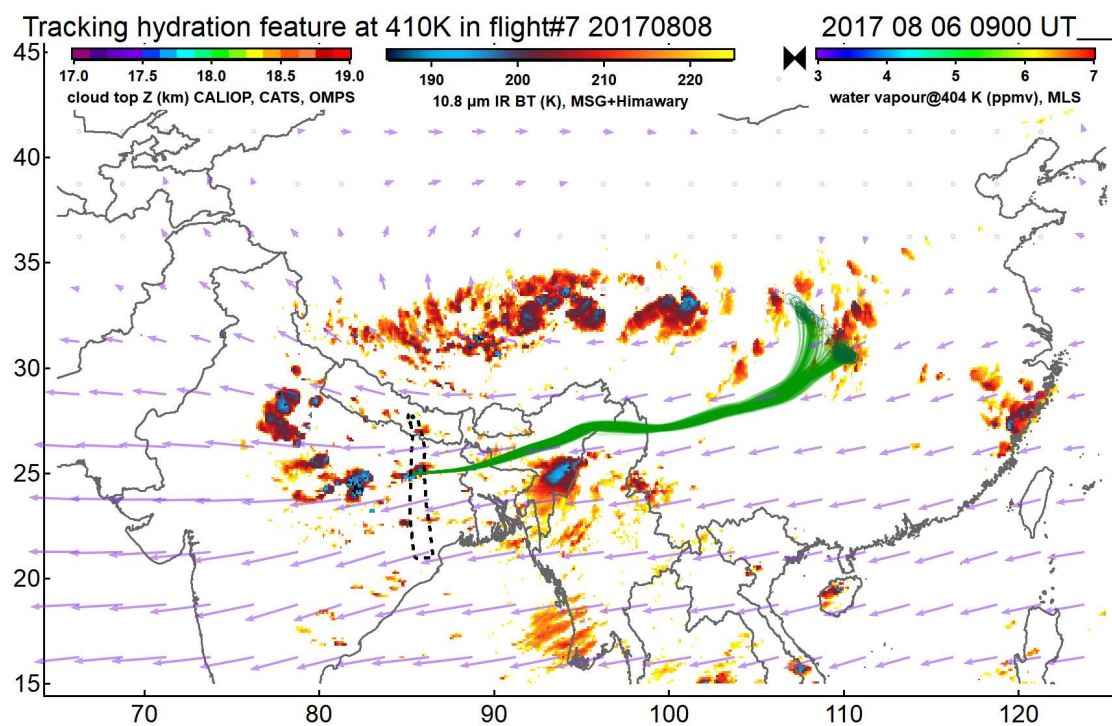


Dynamics is fixed along Lagrangian trajectories *in a variant of microphysical piggybacking.*

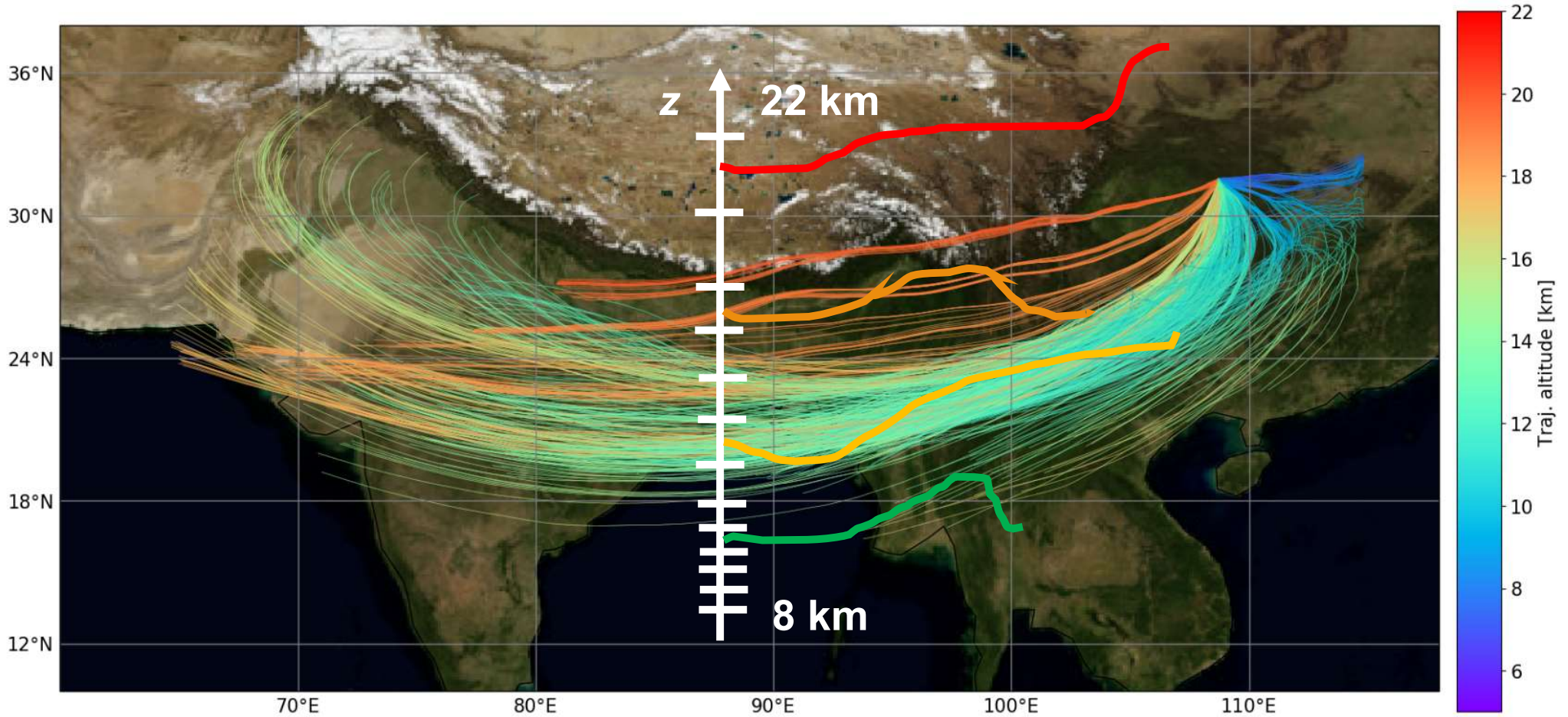


Lagrangian trajectories are initiated in the Sichuan Basin.

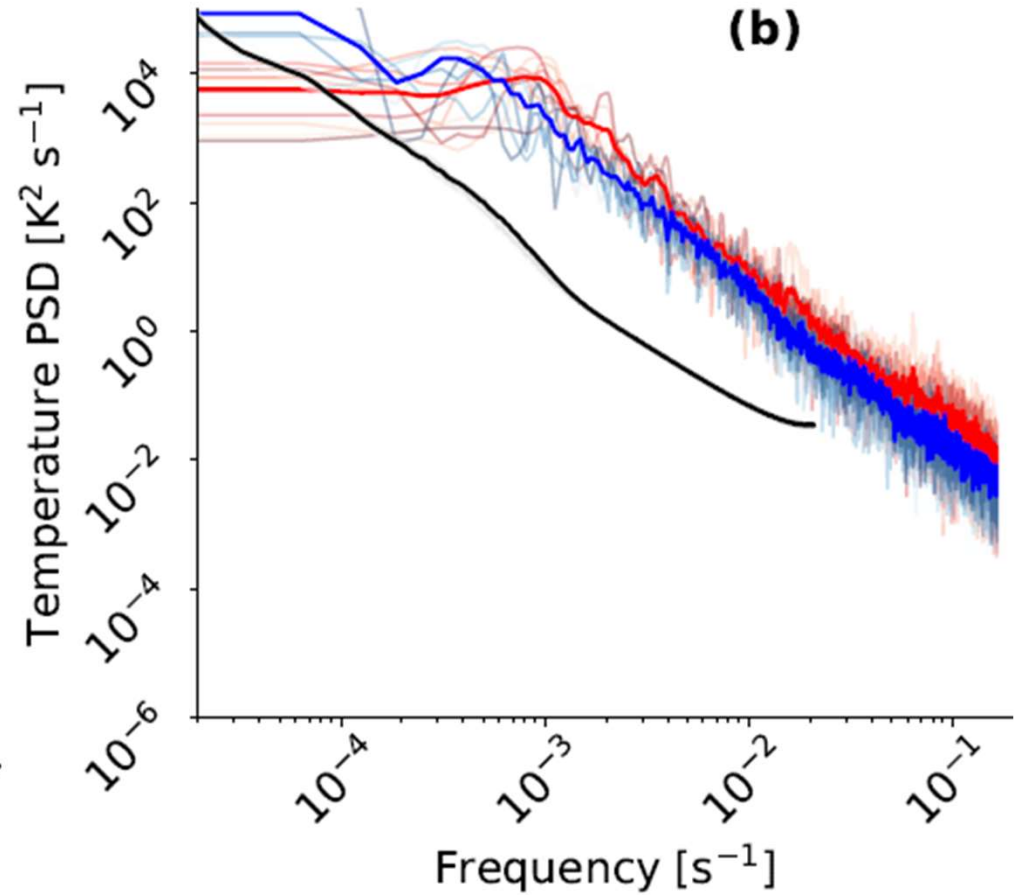
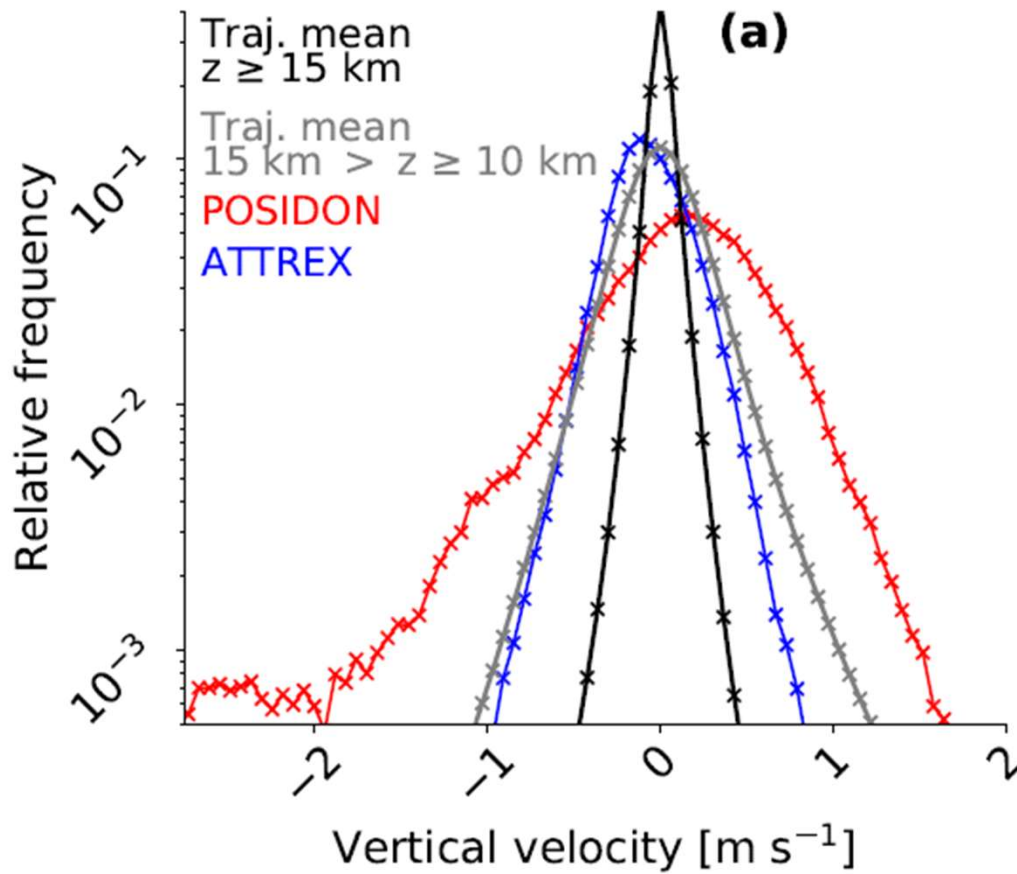
Upper-level moisture during Flight 7 was tracked to convective overshooting in the Sichuan basin.



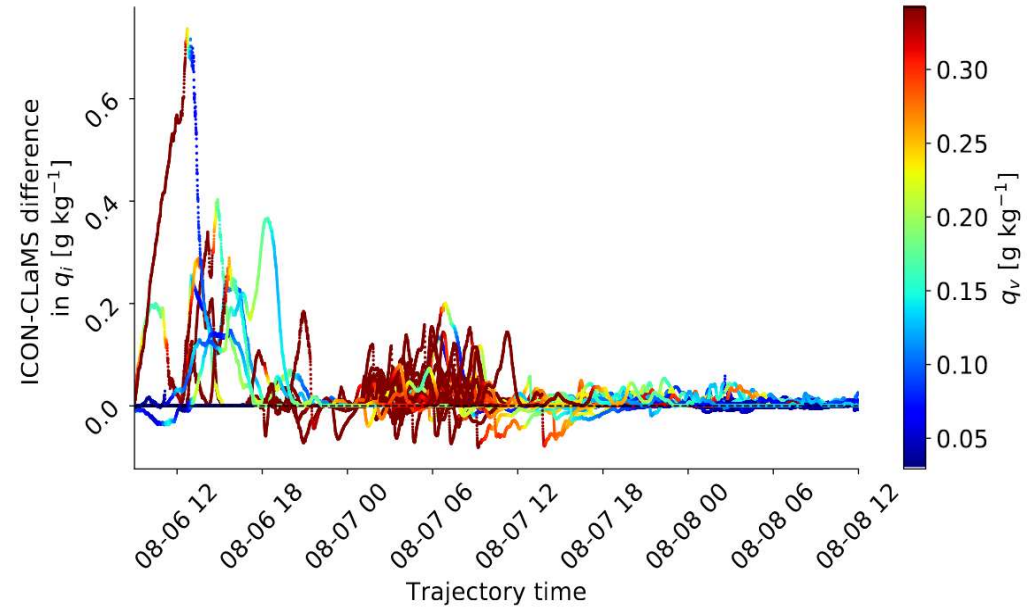
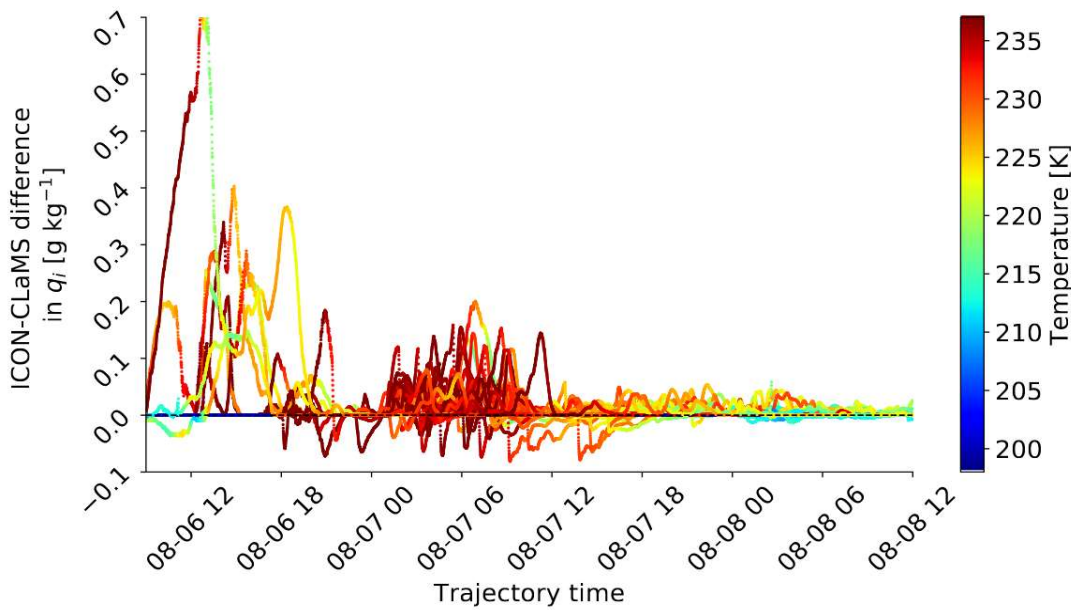
Lagrangian trajectories are initiated in the Sichuan Basin.



What do the trajectory dynamics look like?



The ICON 2-mom scheme predicts much more ice than CLaMS...



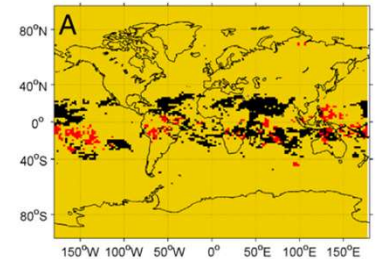
...especially for warmer subzero temperatures and high moisture content.

Summary

Attribution metrics defined from sensitivities identify vertical velocity variability as a crucial input to ice nucleation.

Different regions are driven by sensitivity versus variance.

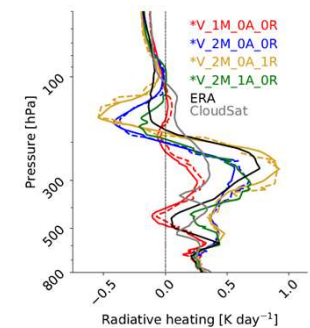
Attributions are model framework-dependent.



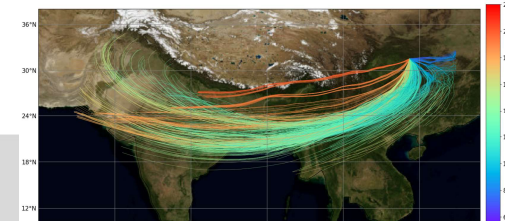
Ice microphysical parameters can change cloud-radiative heating by a factor of 4 and mean OLR by 30 W m^{-2} .

1-mom *versus* 2-mom differences have an “altitudinally-stratified explanation.”

Ice crystal size affects the mass extinction coefficient.

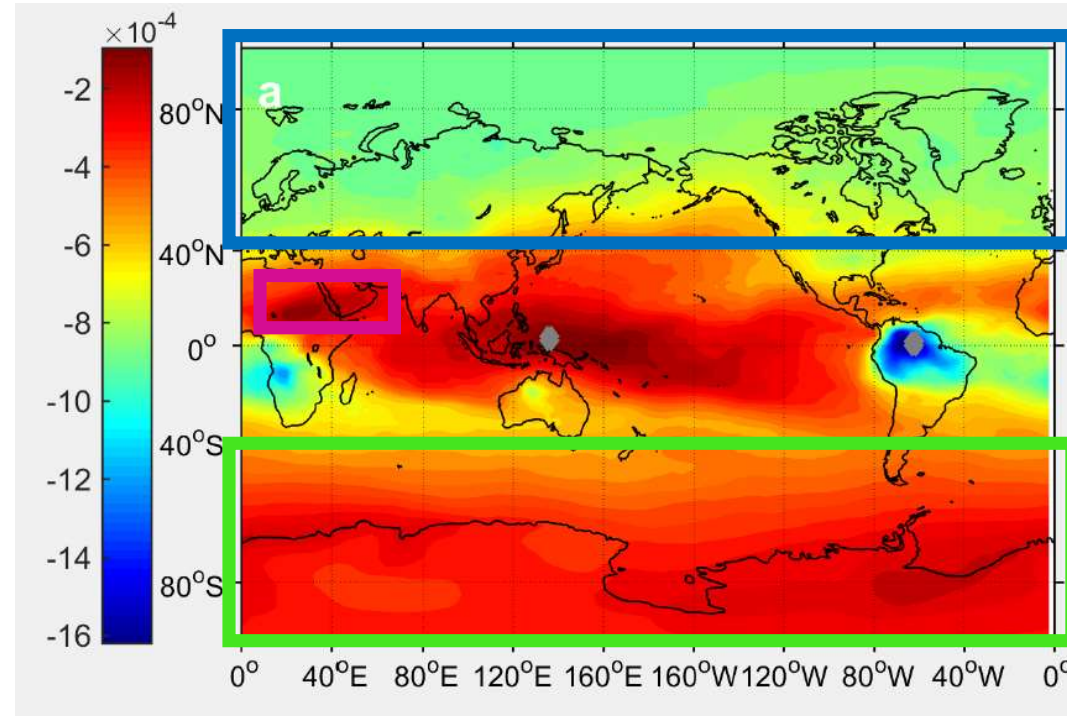
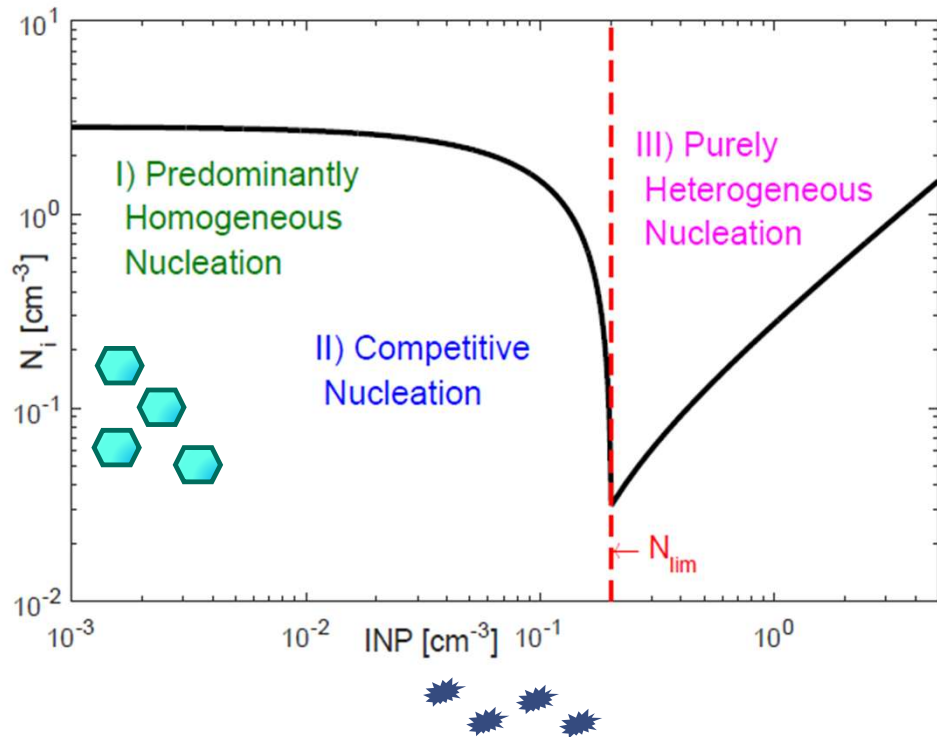


Lagrangian trajectories allow us to isolate the direct impact of microphysics on radiative heating rates.



SUPPLEMENTAL SLIDES

Sensitivities allow us to classify nucleation regime and efficiency.



$$\frac{\partial N_i}{\partial N_{INP}}$$